

# SBND: Short Baseline Near Detector

**Marco Del Tutto**

On behalf of the SBND Collaboration

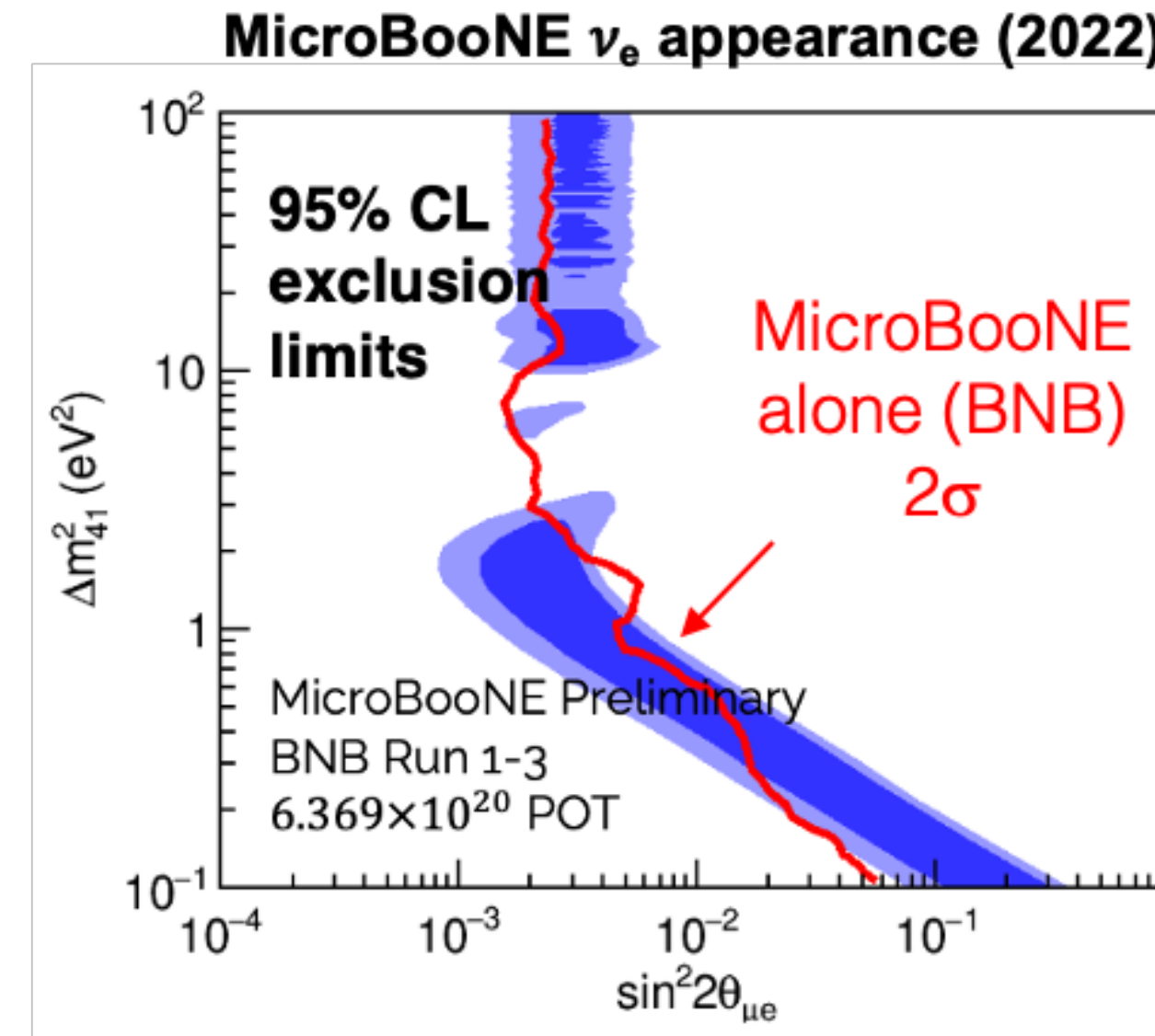
Fermilab Users Meeting

16<sup>th</sup> June 2022

# Introduction

MicroBooNE presented results of their first analyses searching for an excess of low-energy electromagnetic events.

MicroBooNE finds no hints of an electromagnetic event excess but results do not rule out existence of sterile neutrinos.



Entering the next phase of accelerator-based short baseline oscillation searches requires:

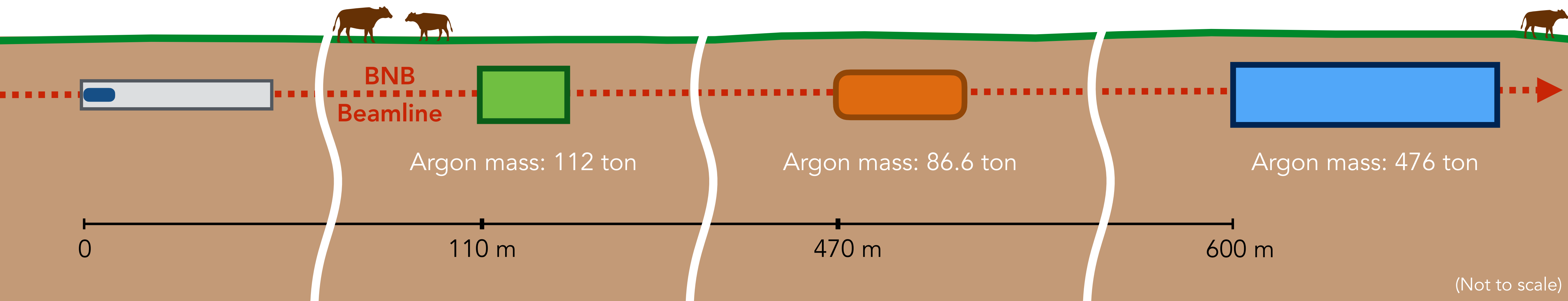
- increased exposure through a larger far detector
- a near detector for systematics constraints

Target

SBND

MicroBooNE

ICARUS



# The Short Baseline Near Detector (SBND)

SBND is the near detector in the Short Baseline Neutrino (SBN) program at Fermilab

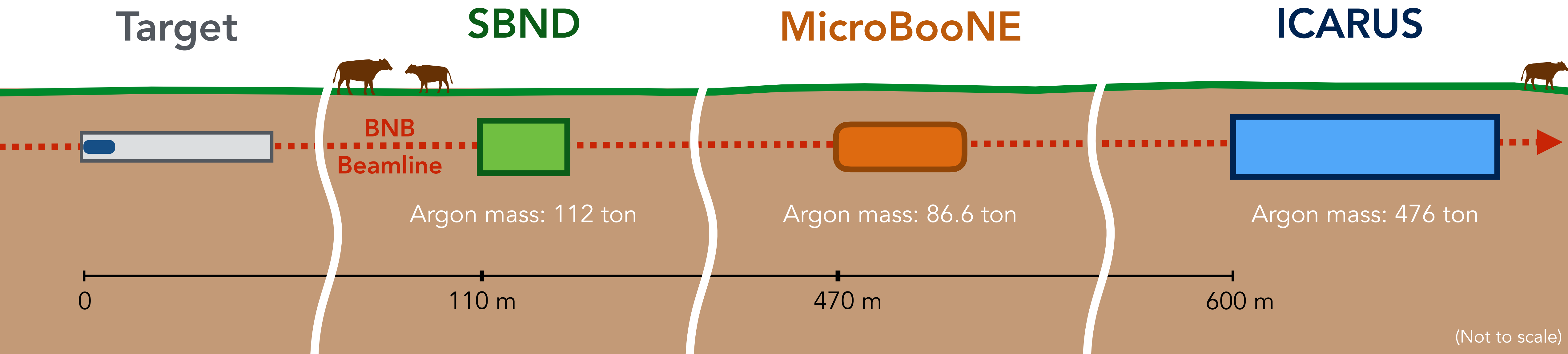
Three Liquid Argon Time Projection Chamber (LArTPC) detectors  
located along the Booster Neutrino Beamline (BNB) at Fermilab

Goals of the SBND:

Search for eV mass-scale sterile neutrinos oscillations

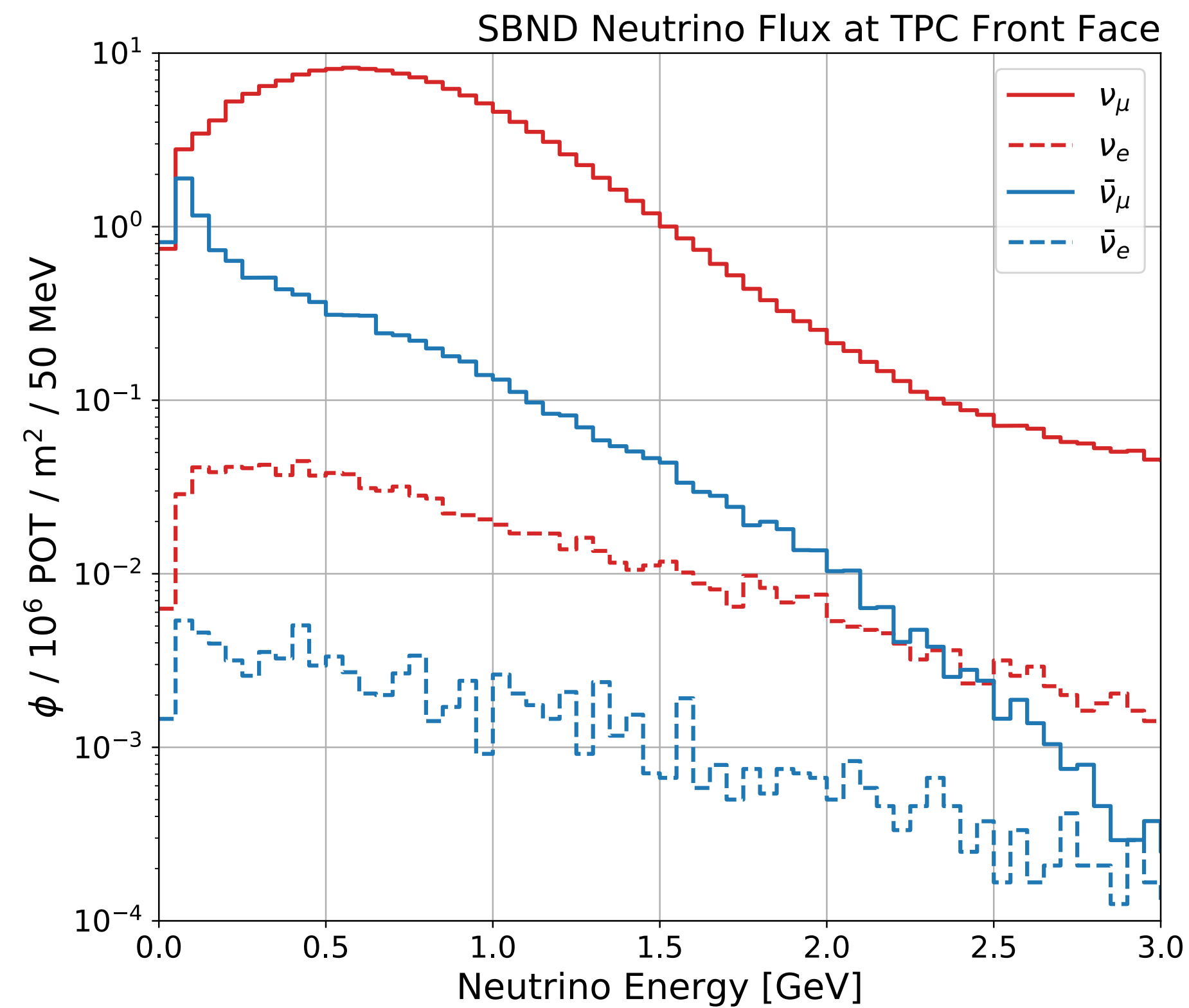
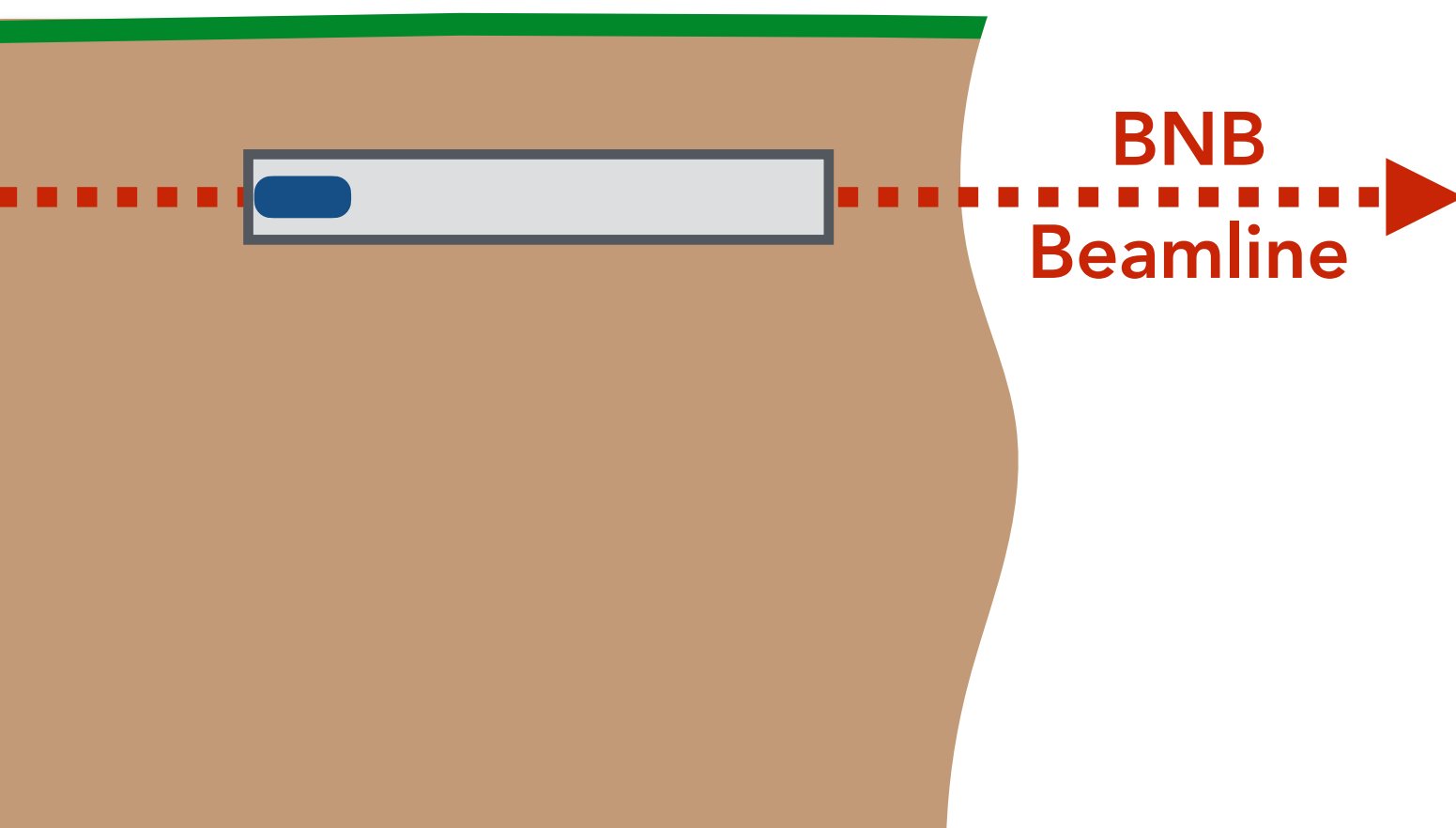
Study of neutrino-argon interactions at the GeV energy scale

Search for new/rare physics processes in the neutrino sector and beyond



# BNB Flux

Target



Neutrino flux at the SBND front face.

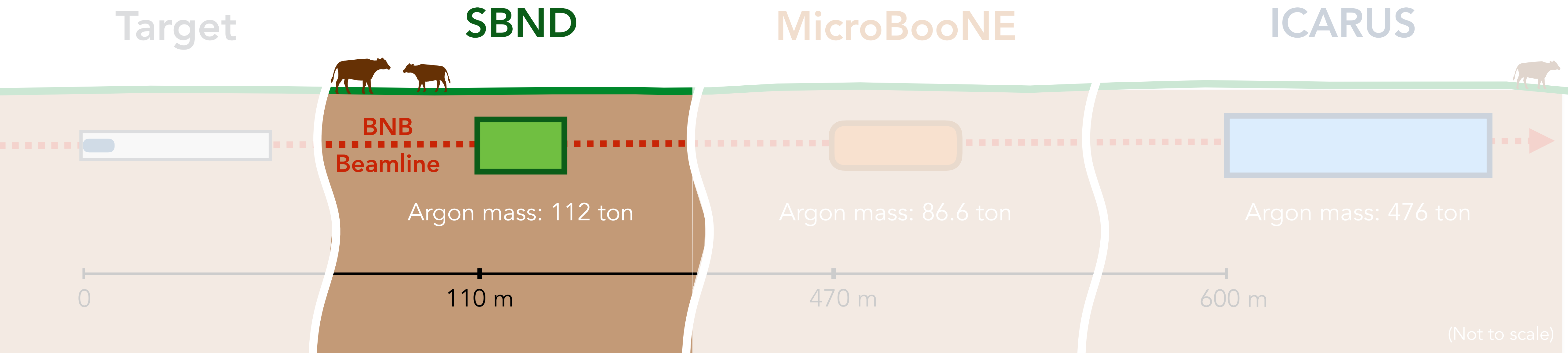
Mean muon-neutrino energy:  $\sim 0.8 \text{ GeV}$

Beam composition:

$\nu_\mu$  (93.6%)

$\bar{\nu}_\mu$  (5.9%)

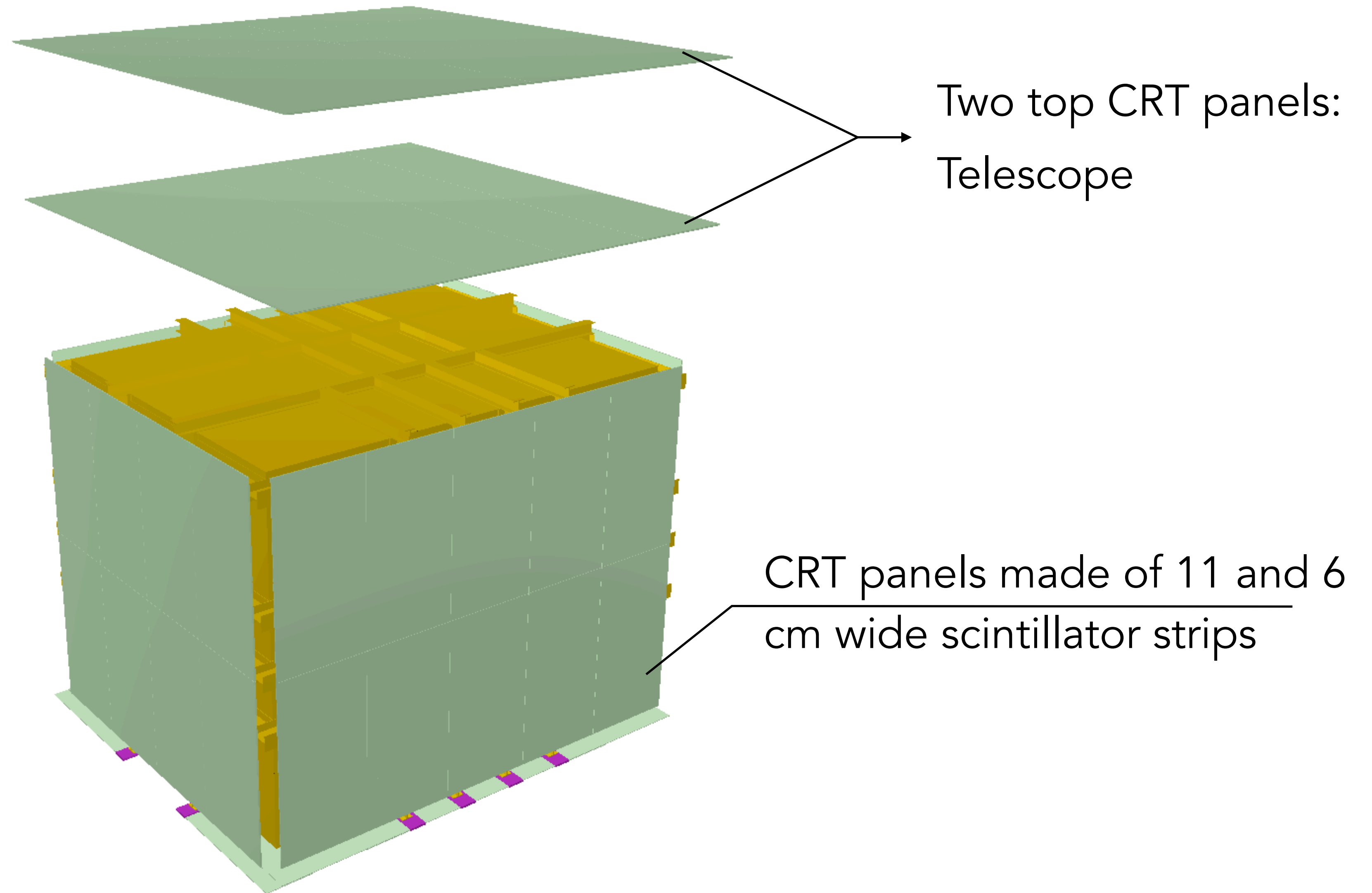
$\nu_e + \bar{\nu}_e$  (0.5%)



# The SBND Detector

Cosmic Ray Tagger  
CRT

SBND will be surrounded by  
scintillator strips to tag  
cosmic rays



# The SBND Detector

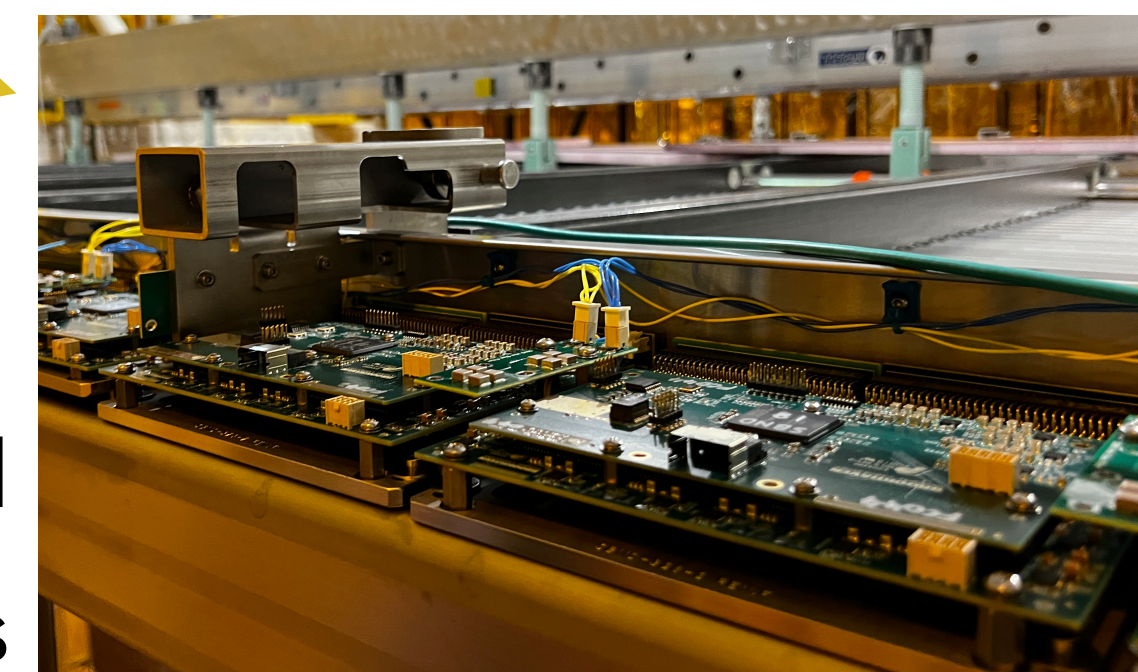
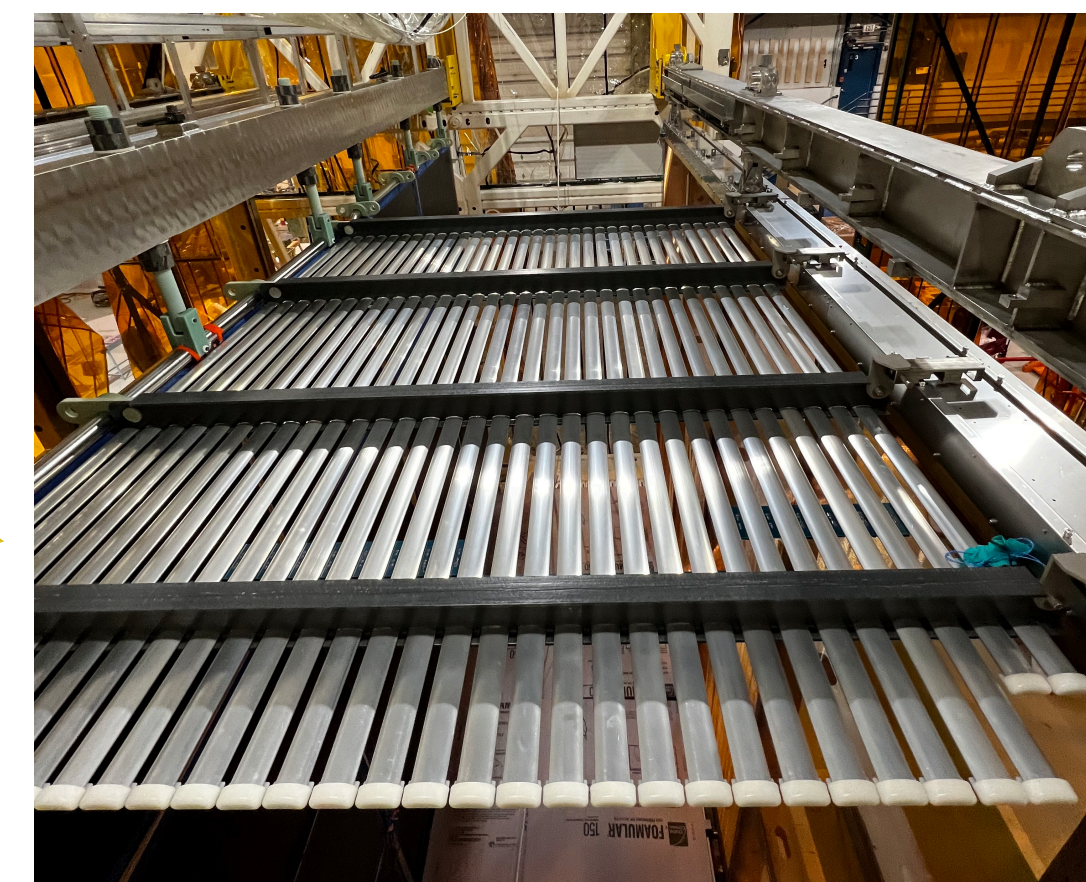
2 Time Projection Chambers  
for a total of 4m x 4m x 5m

Photo Detection System:

120 PMTs  
192 X-Arapucas

Cathode covered with TPB coated reflectors

Field Cage



Cold  
electronics

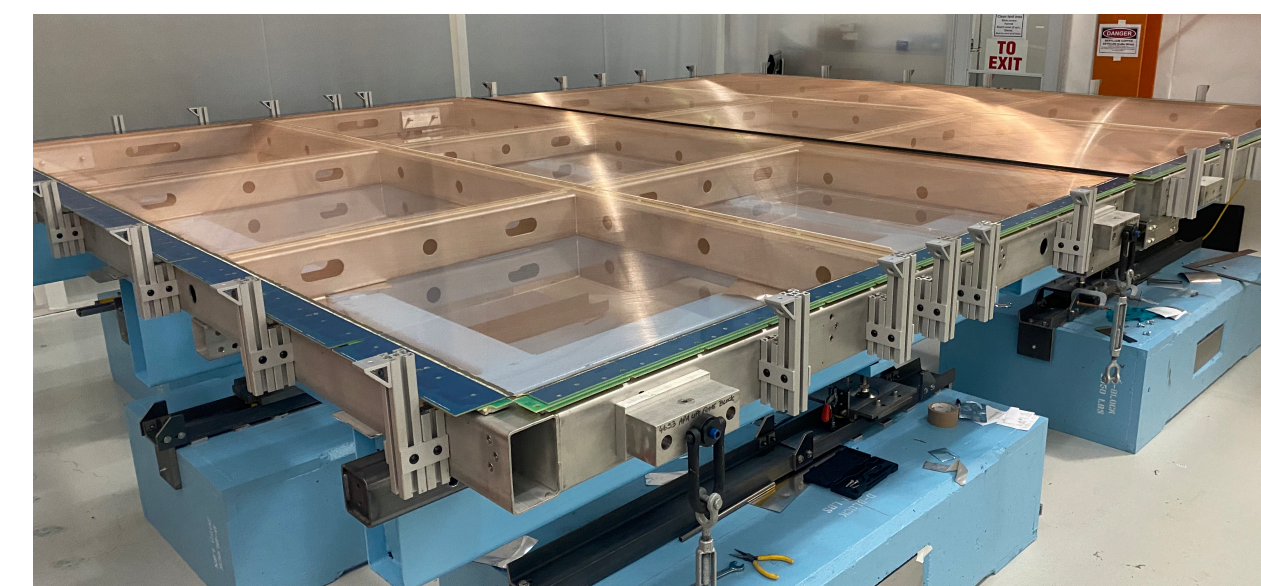
Anode

Cathode

TPC 1

Anode

TPC 0



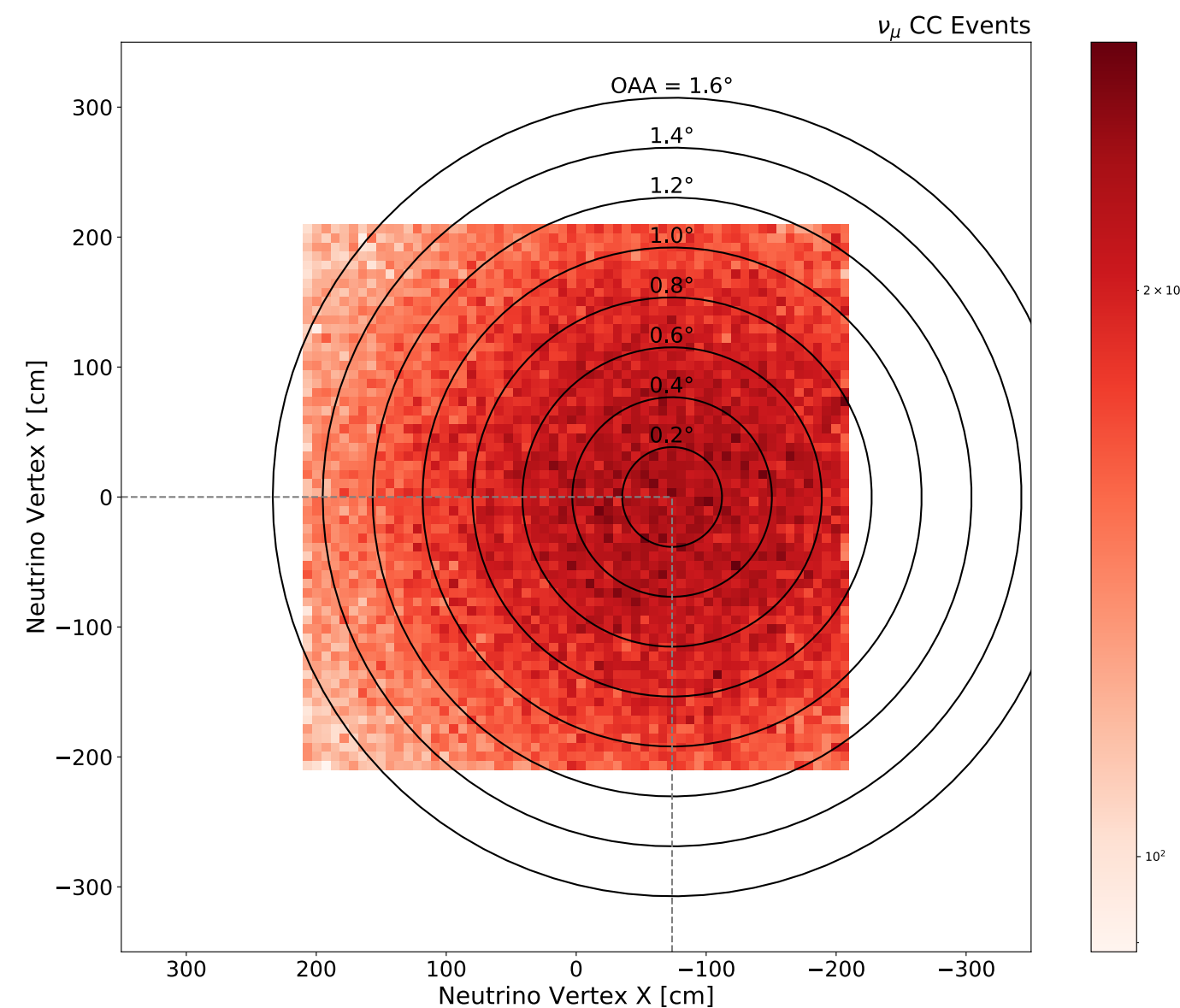
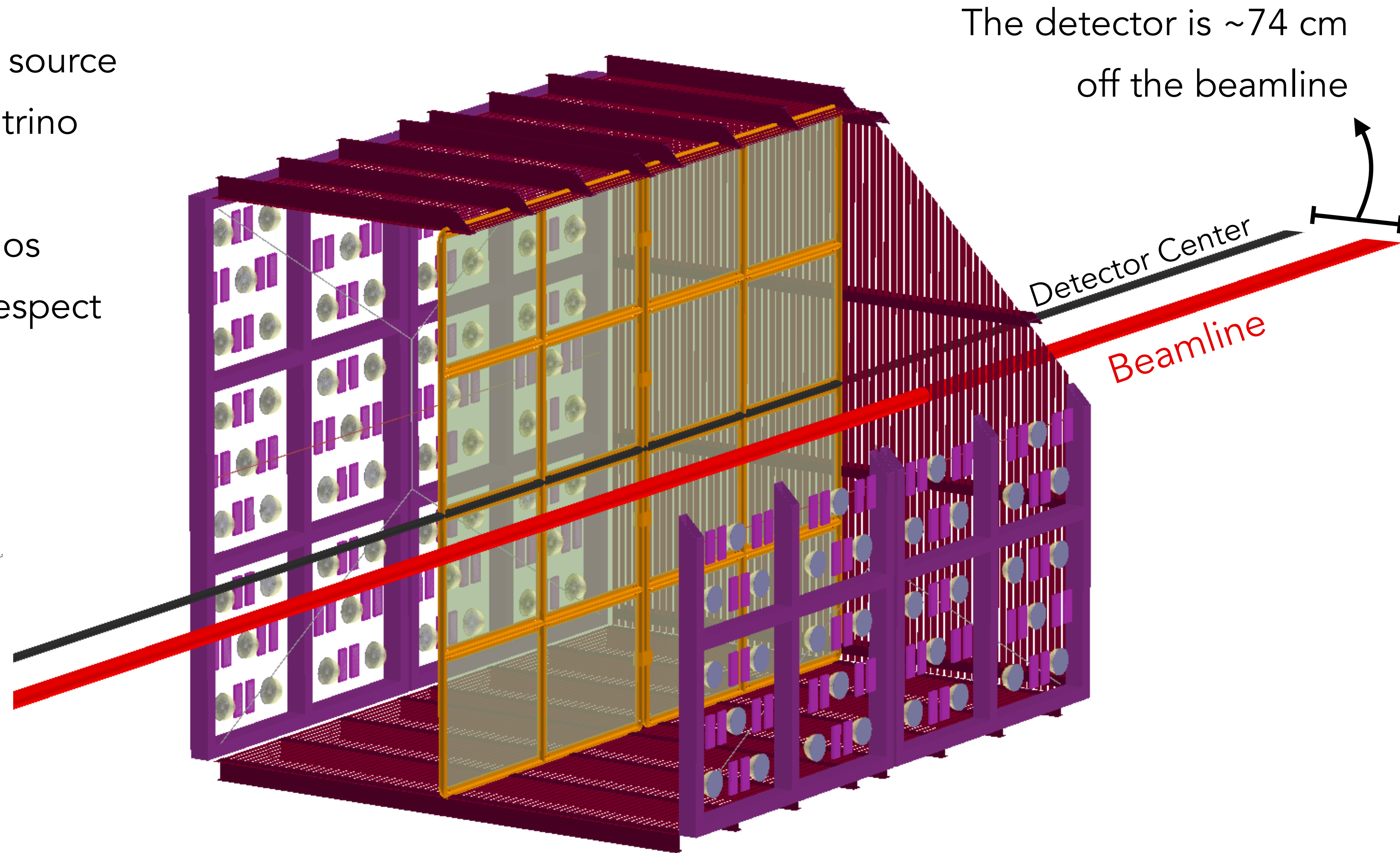
Wire Plane  
3 readout wire  
planes  
~11000 wires

# A Slightly Off-Axis Detector

SBND is:

- very close (110 m) to the neutrino source
- not perfectly aligned with the neutrino beamline

The detector is traversed by neutrinos coming from different angles with respect to the beam axis.



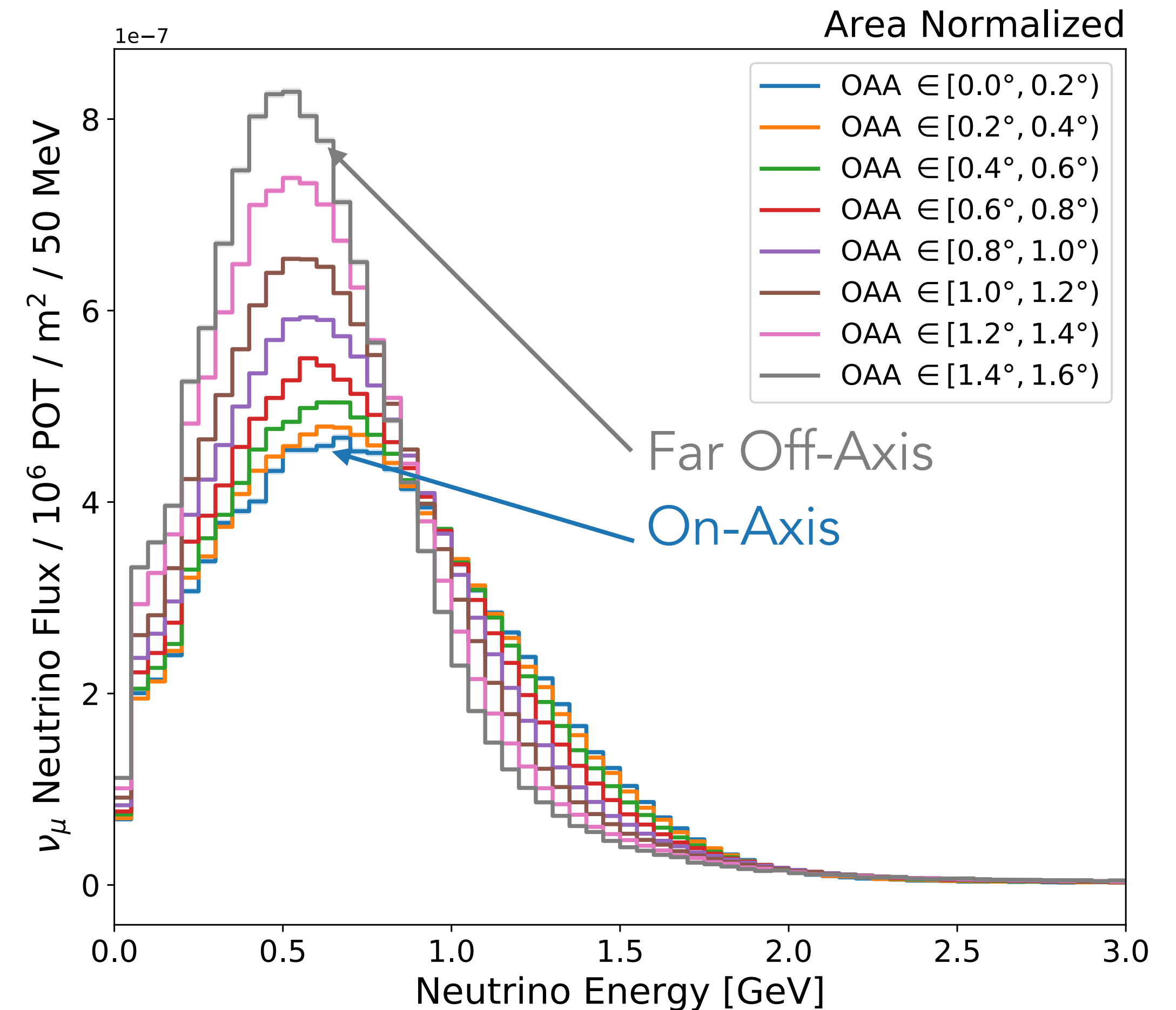
# SBND-PRISM Fluxes

This “PRISM” feature of SBND allows sampling of multiple neutrino fluxes using the same SBND detector.

Similar to the nu-PRISM and DUNE-PRISM concepts, but with a fixed detector.

The neutrino energy distributions are affected by the off-axis position

Muon neutrino flux in each of the OAA regions



# SBND Construction

- SBND TPC construction is now completed (as of this week!)
- Wire planes, field cage, and cold electronics are already installed
- Photon detection system is being installed now



TPC



TPC



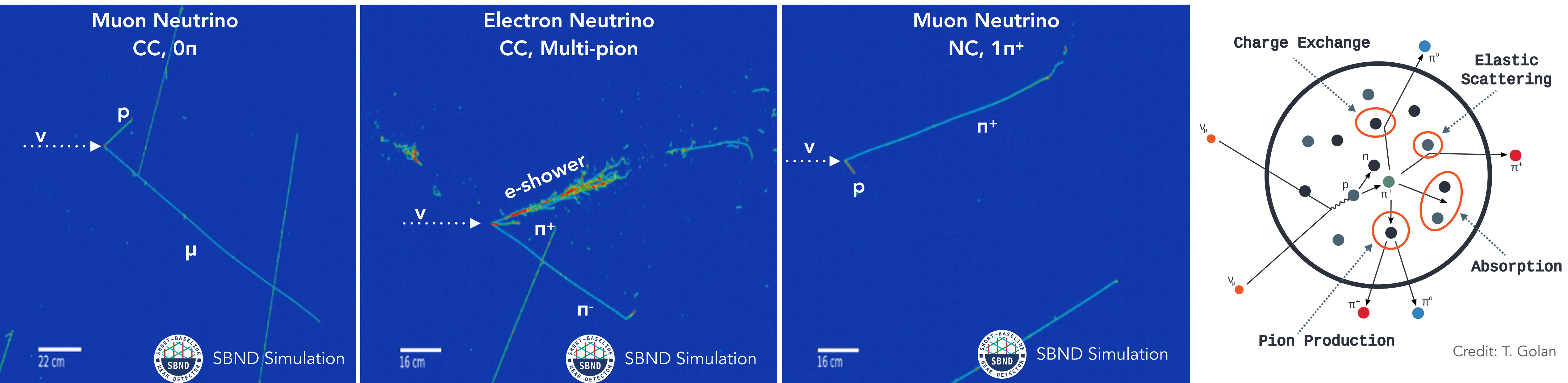
Cryostat construction is in progress

# SBND Physics Program

- Neutrino cross-section measurements
- Search for eV mass-scale sterile neutrinos oscillations in the SBN program
- Search for new and exotic physics signals

# Cross-Section Measurements

- Understanding neutrino interactions with argon is crucial for the success of current and future neutrino experiments.
- Argon has a complex and heavy nucleus and not much data exists to constrain current generators.
- MicroBooNE measurements provide invaluable insights and SBND will be able to expand on those with an incredible large-statistic dataset allowing to make multi-dimensional analyses with high statistics.
- SBND measurements will largely improve DUNE physics as SBND's kinematics cover large parts of DUNE's phase space.



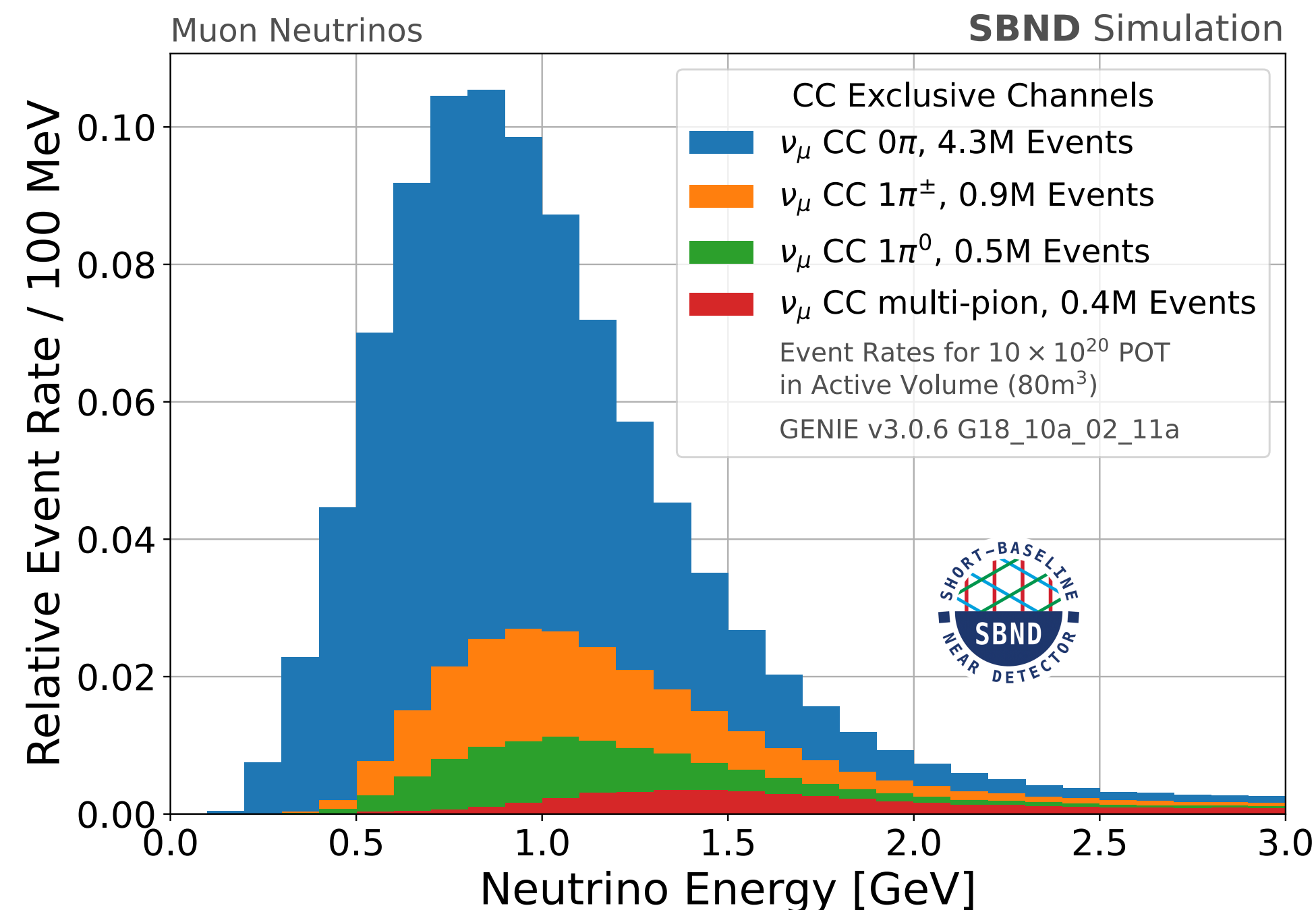
# Cross-Section Measurements

SBND data will enable a generational advance in the study of neutrino-argon interactions in the GeV energy range, with low thresholds for particle tracking and calorimetry and enormous statistics.

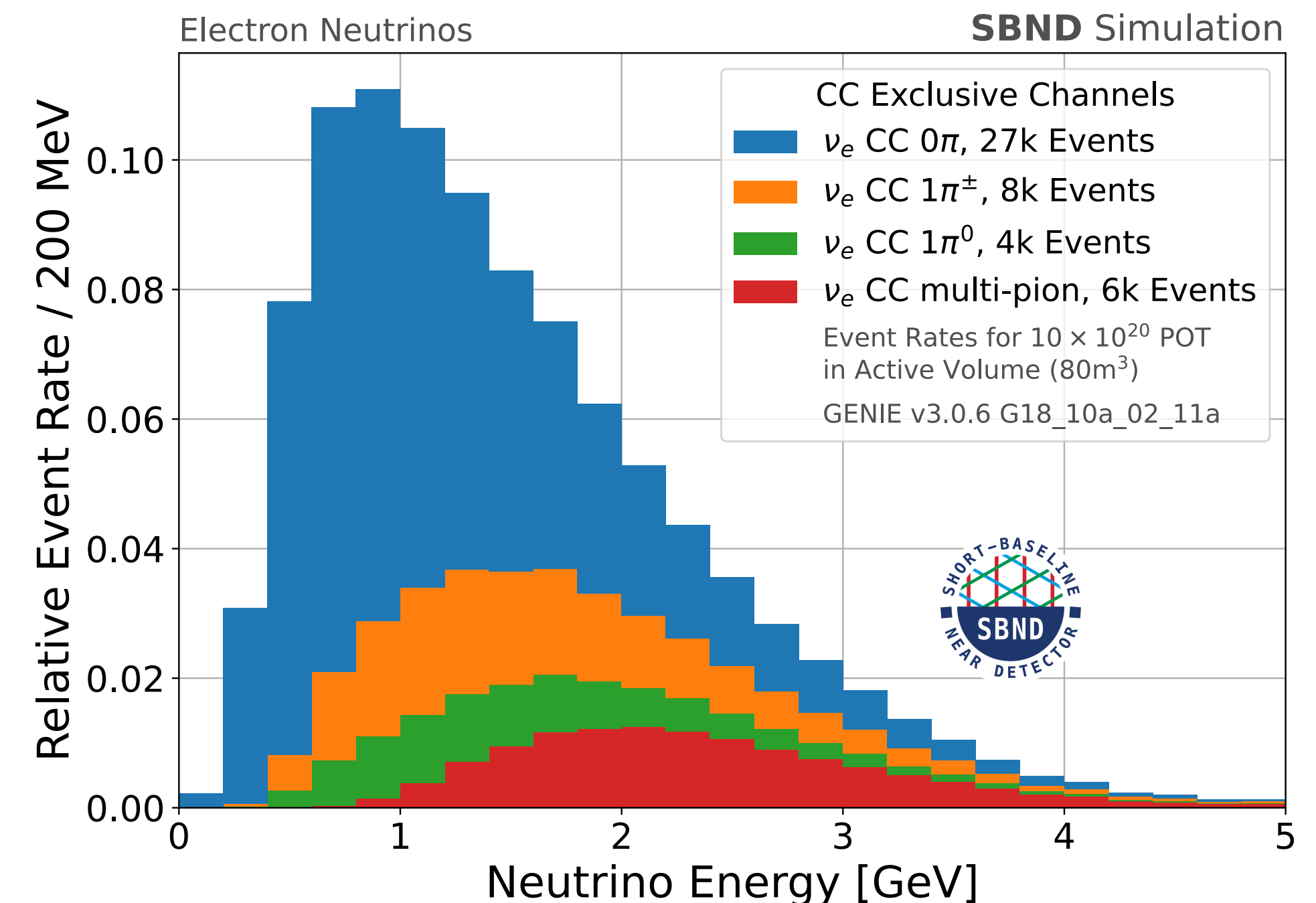
SBND will have the largest dataset of  $\nu$ -Ar interactions and will do high-statistics measurements of many signatures and can observe rare channels.

SBND will record 20-30x more neutrino-argon interactions than is currently available.

**SBND will observe 5000  $\nu$ -events/day!**

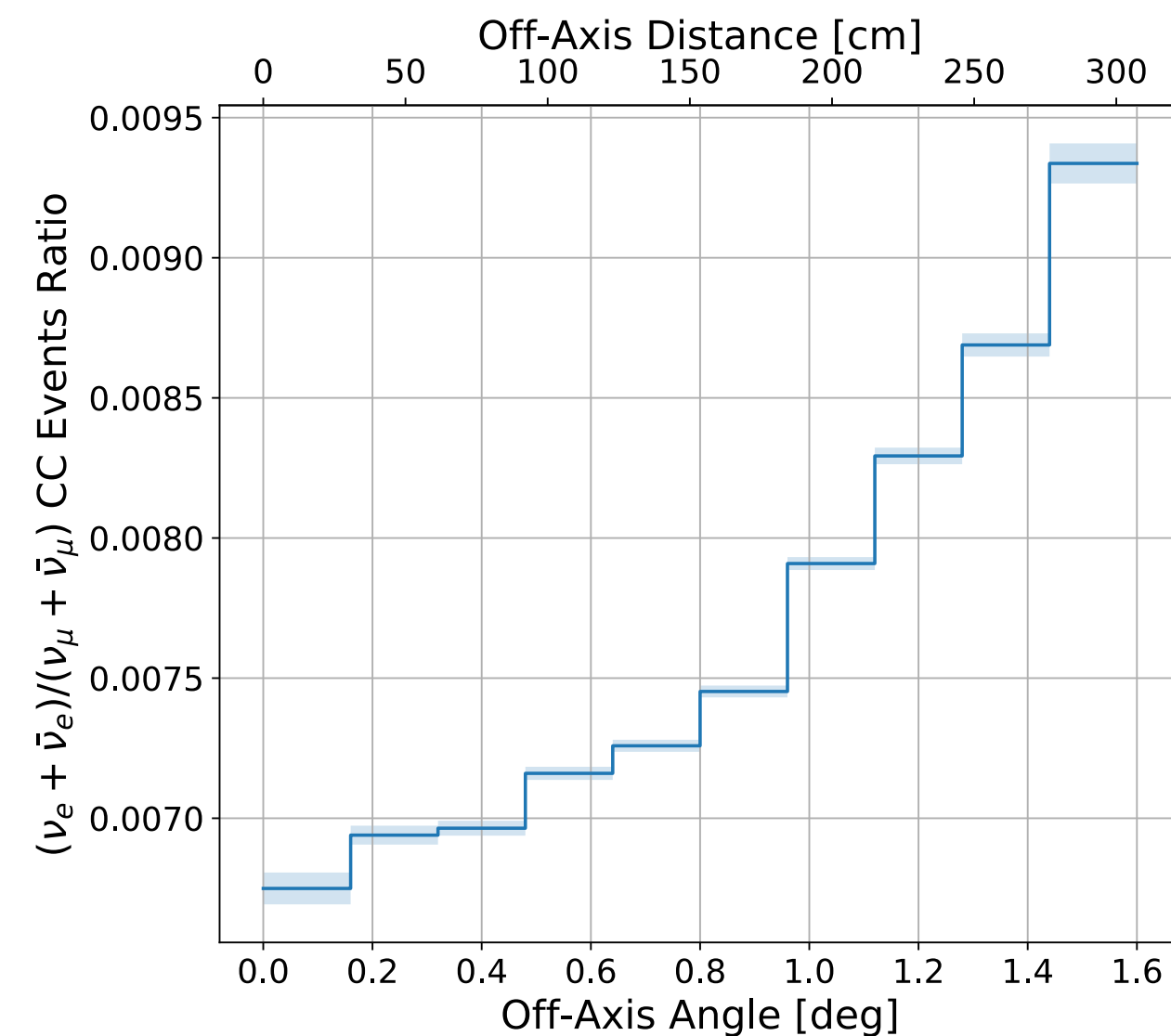
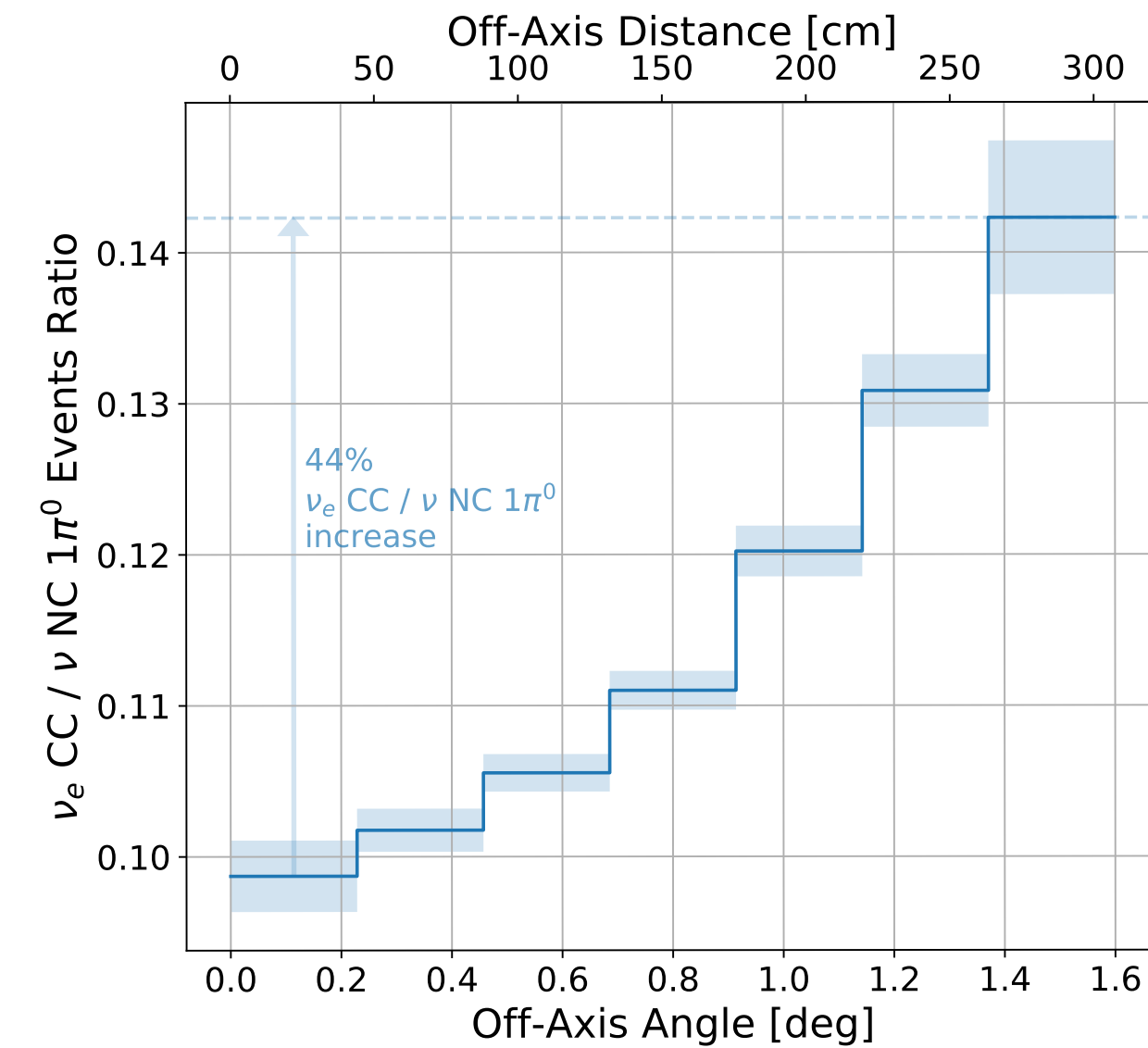
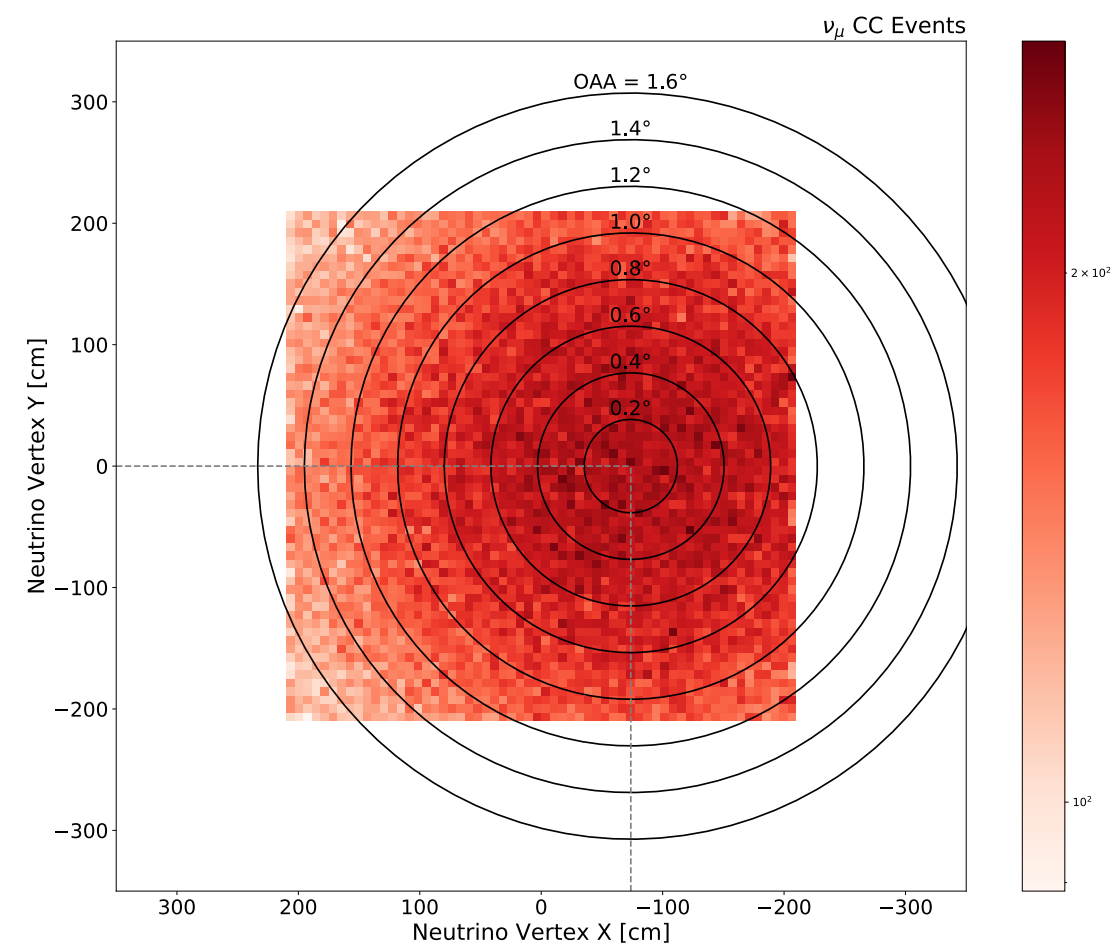


1.5M  $\nu_\mu$  CC events in 1 year



12k  $\nu_e$  CC events in 1 year

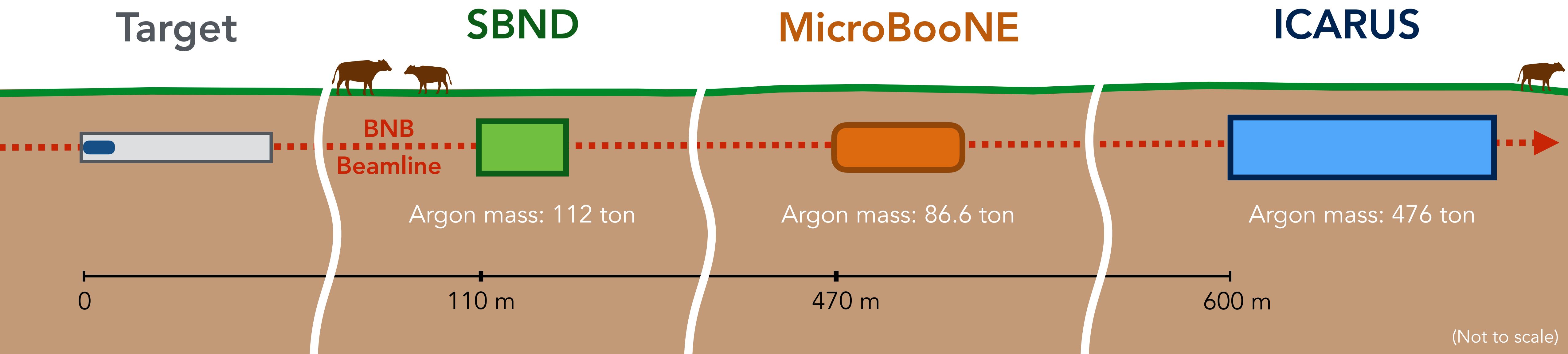
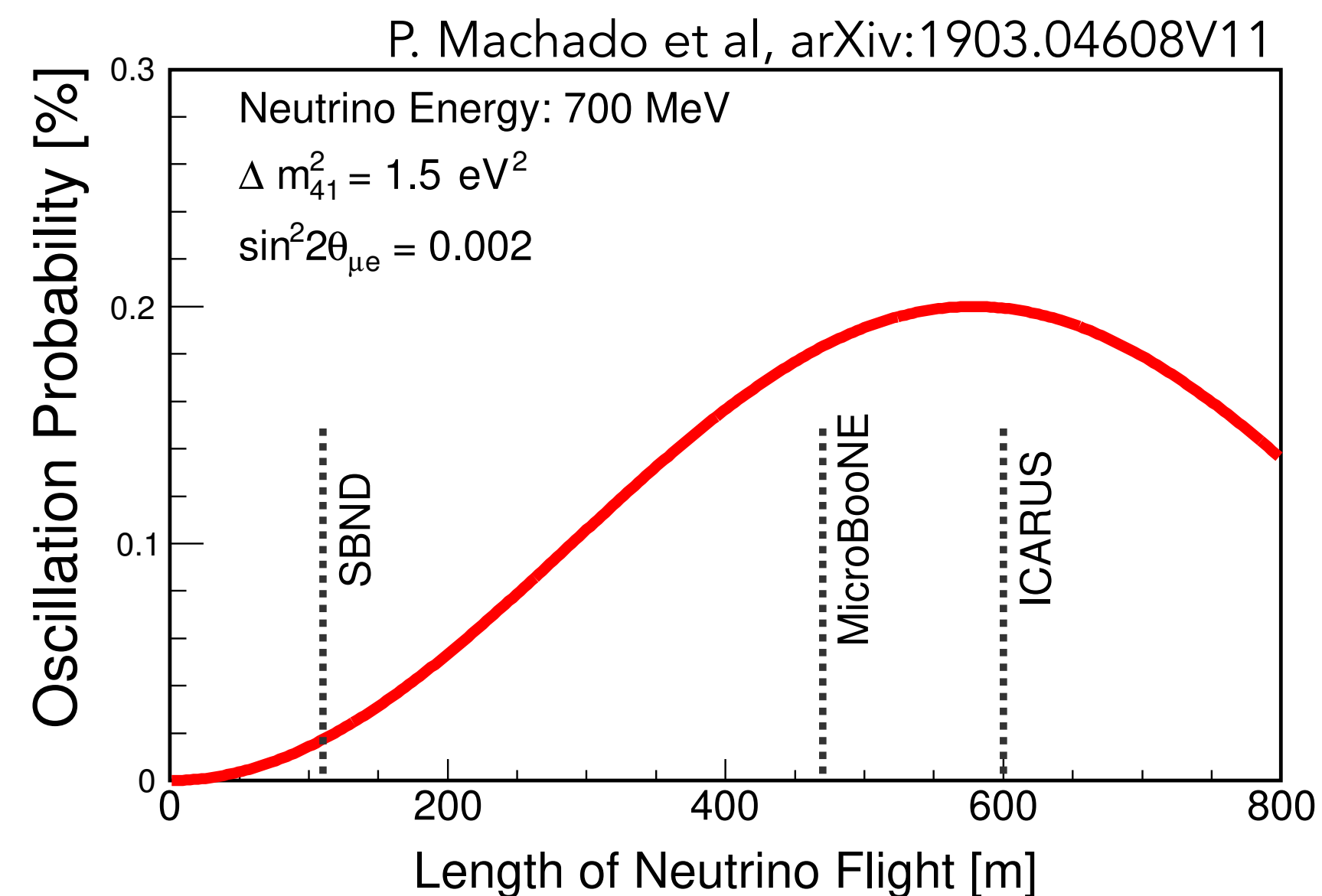
# Cross-Section Measurements with SBND-PRISM



- PRISM provides a natural way to reduce background by moving off-axis.
- Note that we expect high event statistics in all off-axis regions.
- Going off-axis, the increase in  $\nu_e$  to  $\nu_\mu$  flux ratio combined with a choice of kinematics where  $\nu_e$  to  $\nu_\mu$  differences are prominent should allow us to measure the  $\nu_e/\nu_\mu$  cross section (can study lepton mass effects).

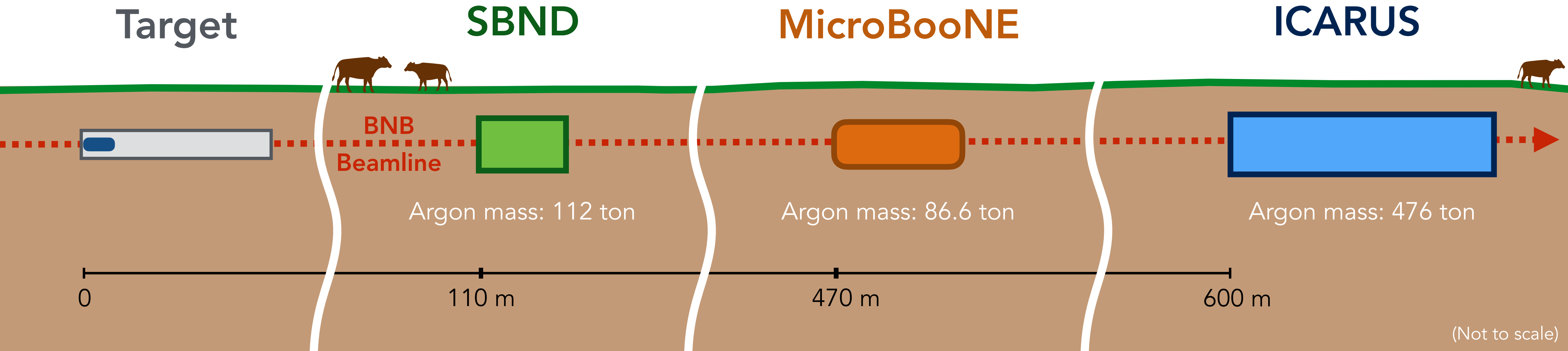
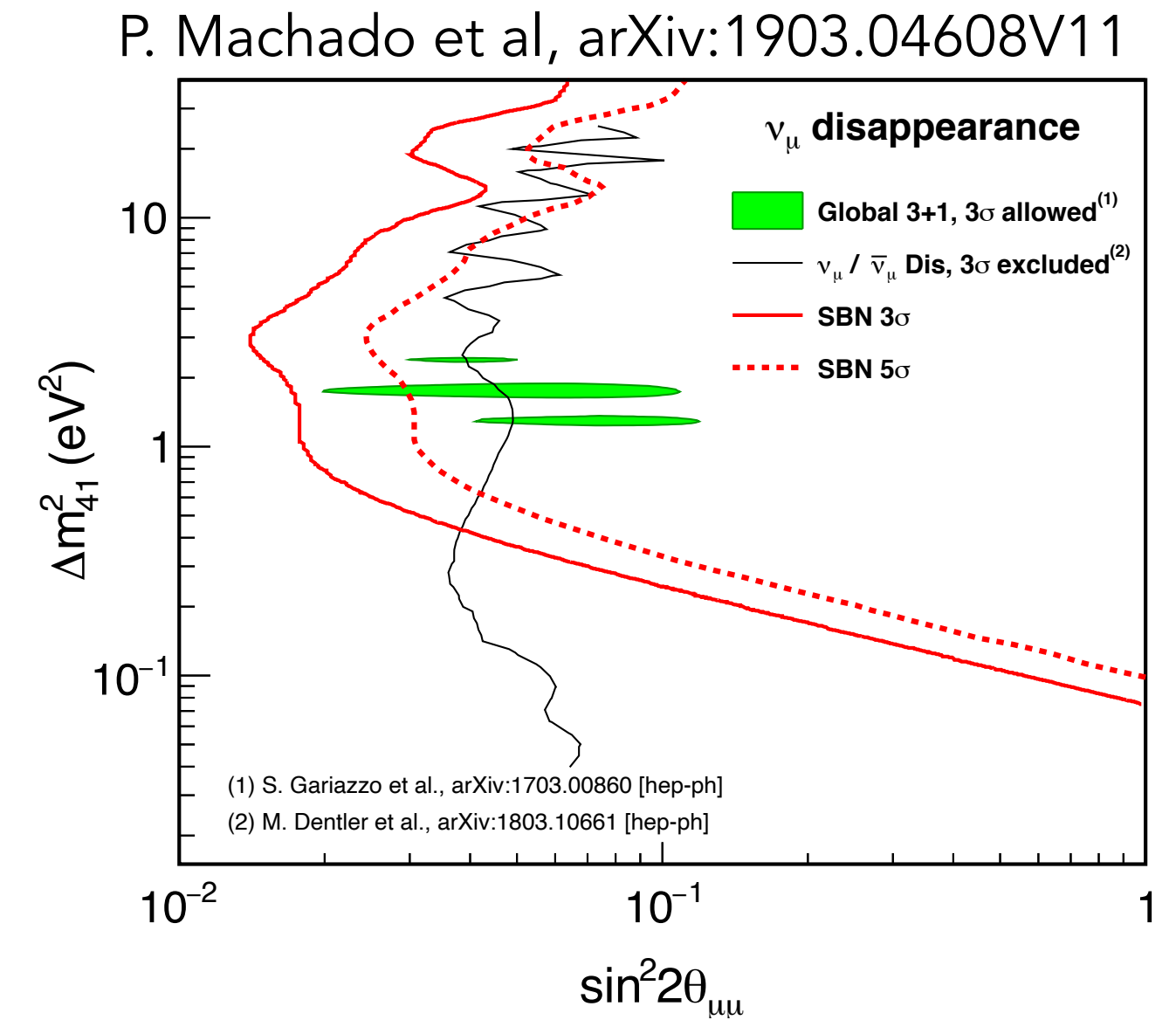
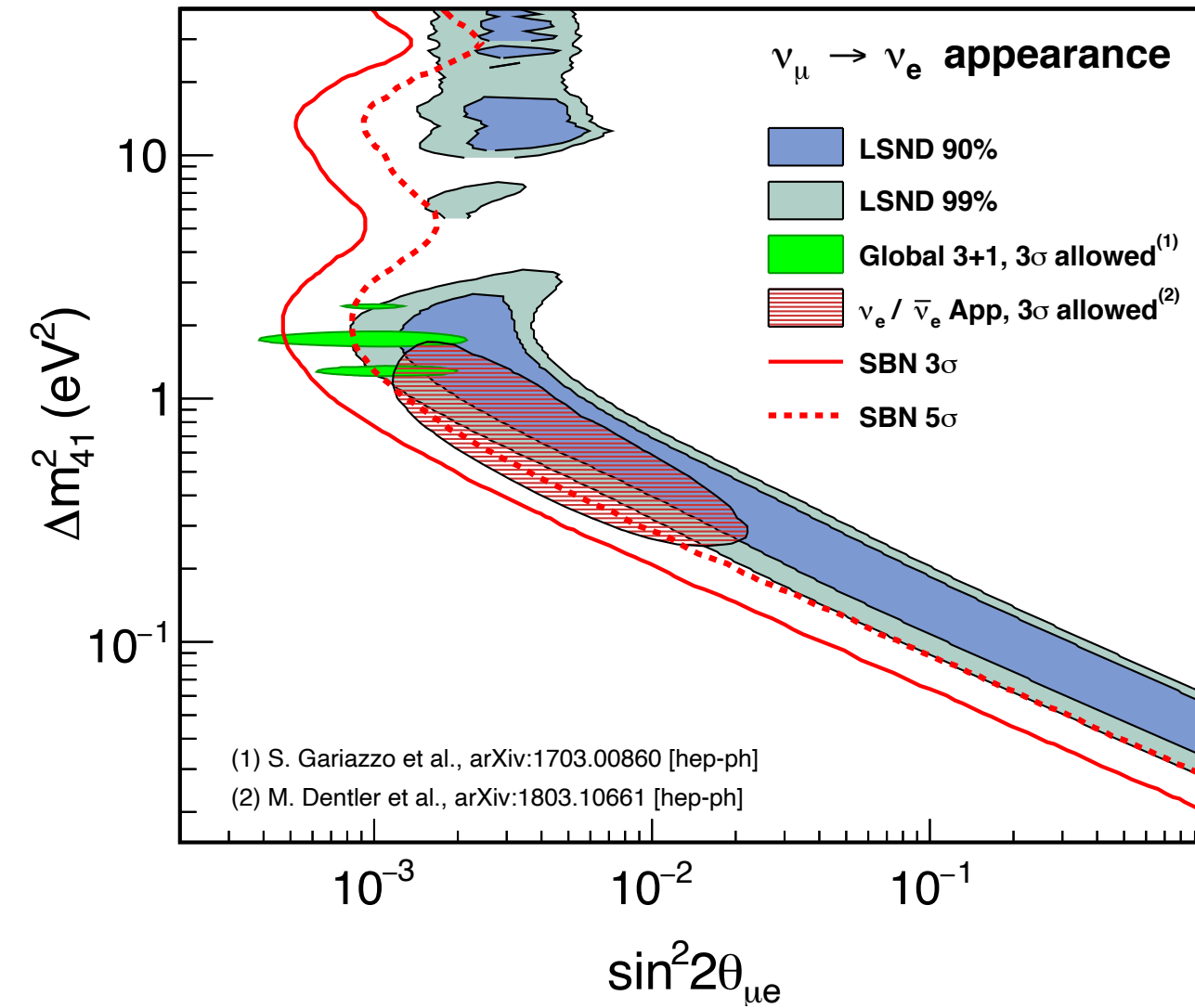
# SBN Oscillation Sensitivity

- SBND + ICARUS will test the sterile neutrino hypothesis
- Three detectors sampling the same neutrino beam at different distances.
- Effective systematics constraint through the near detector (SBND) and same detector technology in the near and far detector.



# SBN Oscillation Sensitivity

- Search for both  $\nu_e$  appearance and  $\nu_\mu$  disappearance
- SBND + ICARUS will test the sterile neutrino hypothesis: can cover the parameter space favored by past anomalies with  $5\sigma$  significance



# Alternative Explanations

There is tension when combining  $\nu_e$  appearance and  $\nu_\mu$  disappearance data.

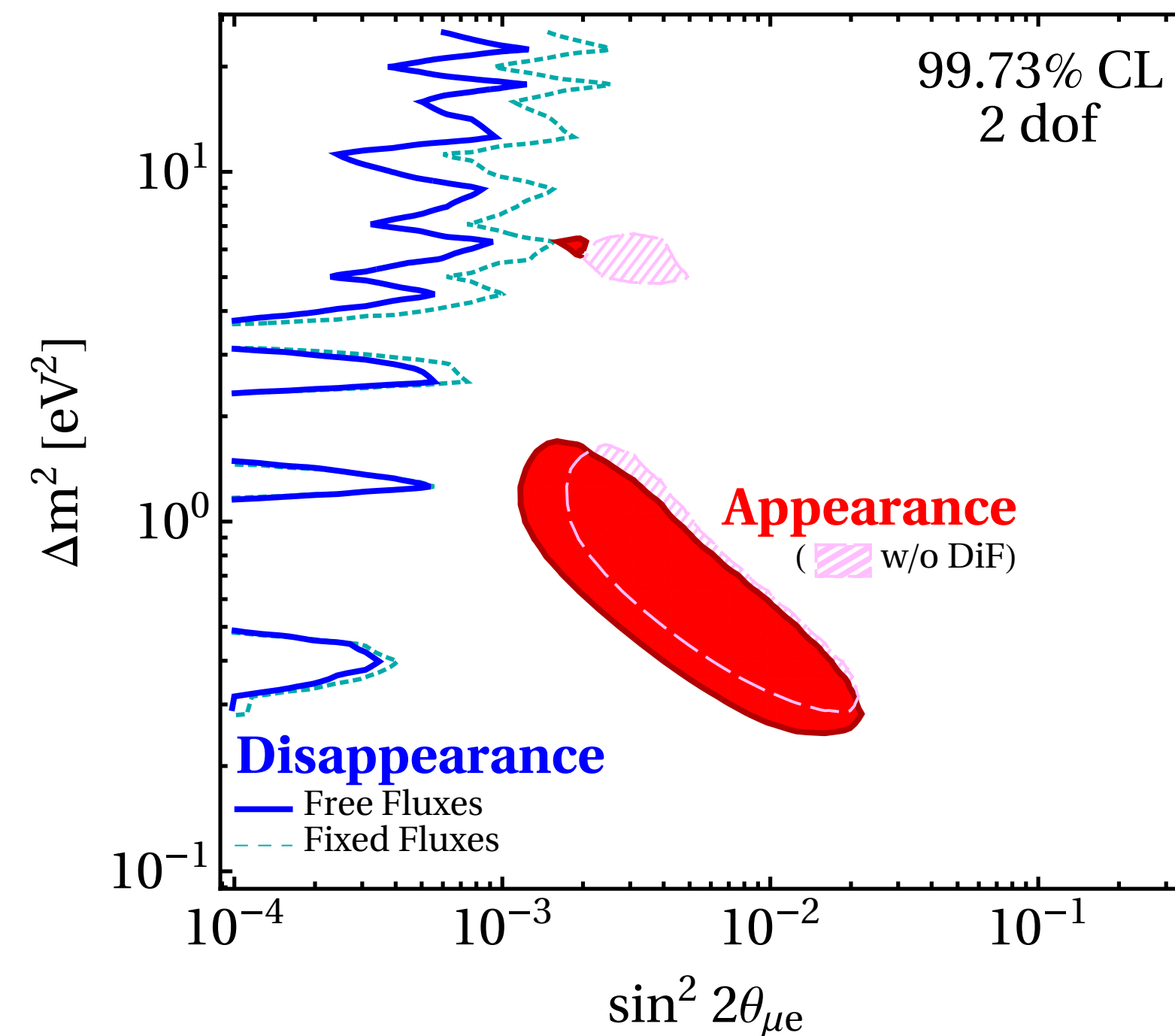
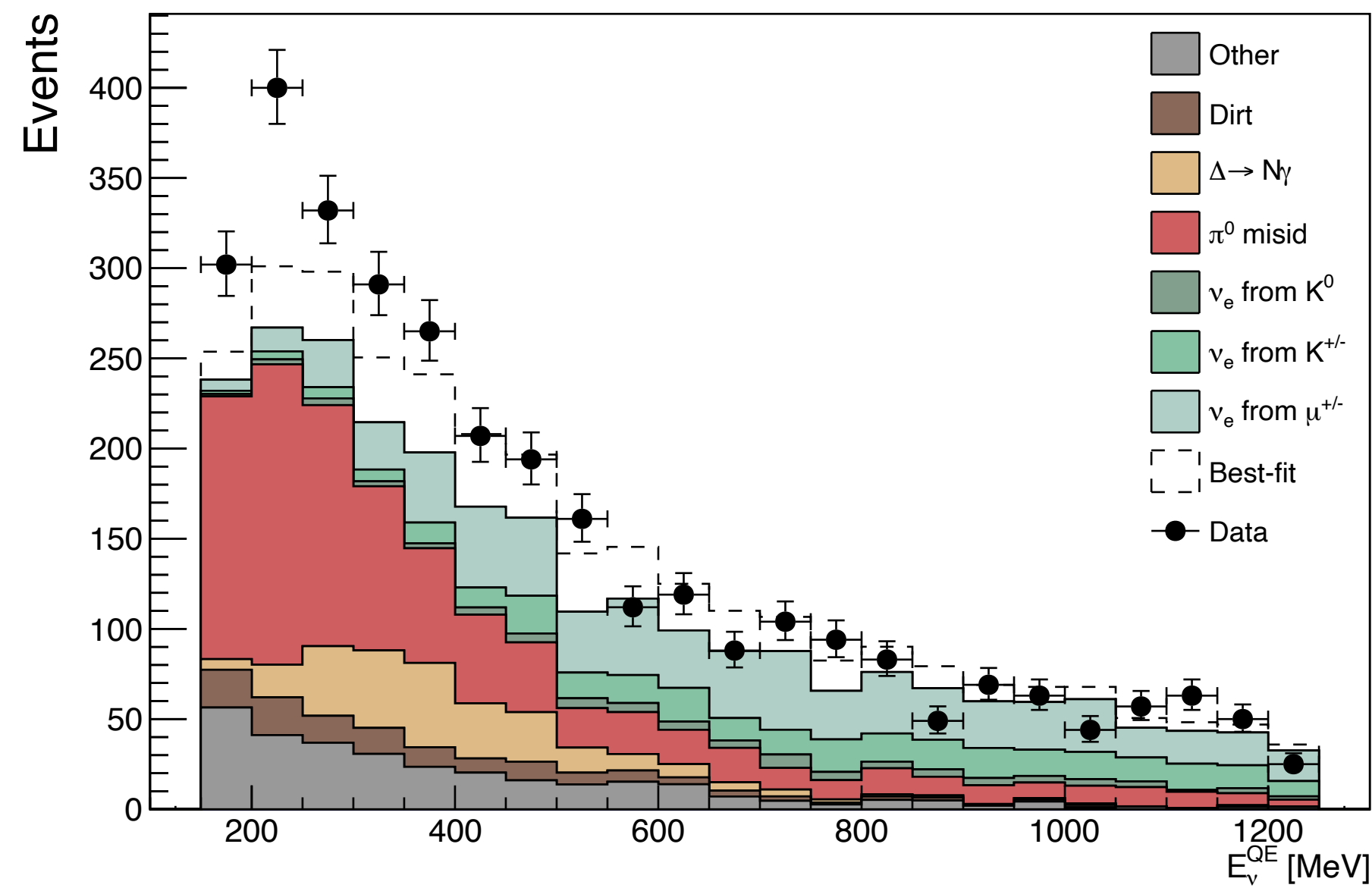
This tension excludes a sterile neutrino

oscillation explanation of the  $\nu_\mu \rightarrow \nu_e$  anomalies at the  $4.7\sigma$  level.

Alternative explanations exist that could explain the MiniBooNE excess.

MiniBooNE  
electron-like  
excess

Phys. Rev. D 103,  
052002 (2021)



Limits from the  
disappearance and  
allowed appearance  
regions

M. Dentler et al., JHEP  
08:010 (2018)

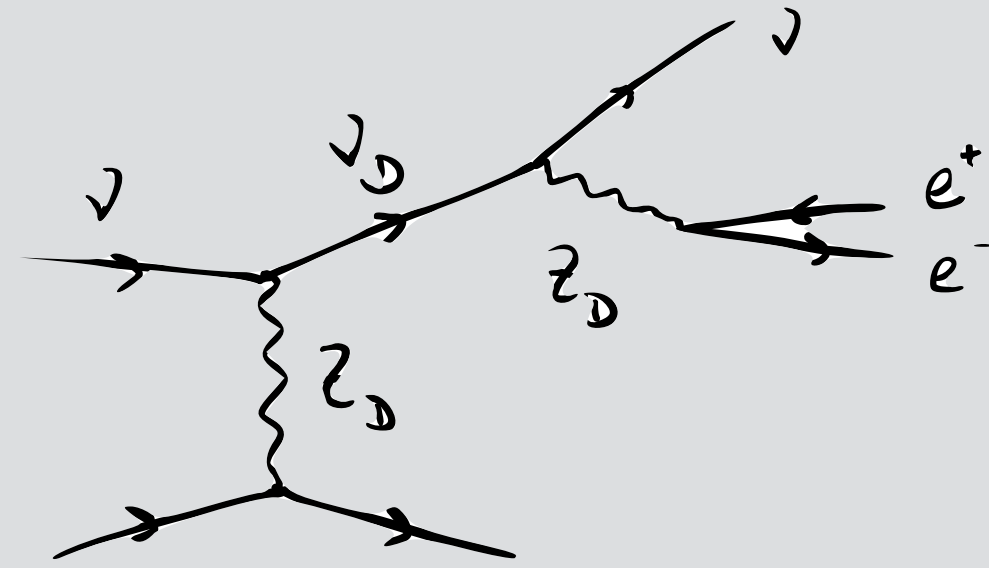
# Beyond Standard Model Searches

Alternative explanations  
to the MiniBooNE excess  
and other BSM scenarios

Not an exhaustive list

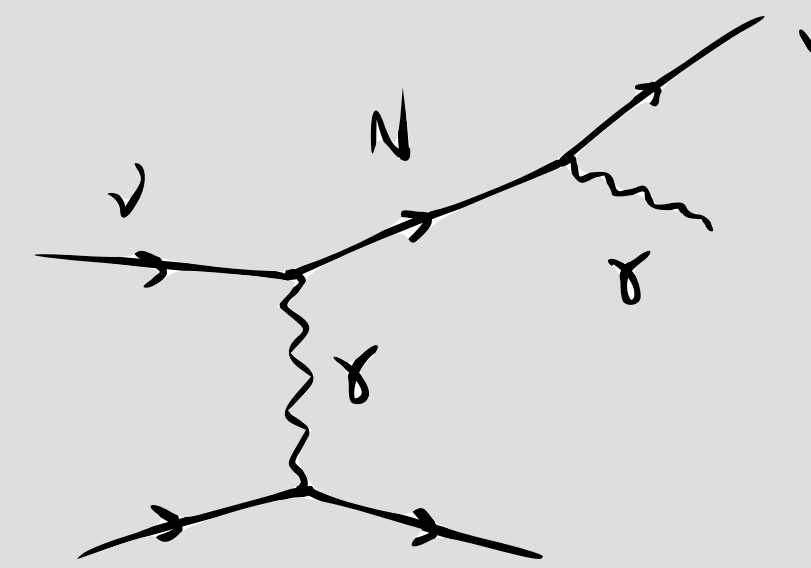
Some diagram credit: Pedro Machado

## Dark Neutrinos



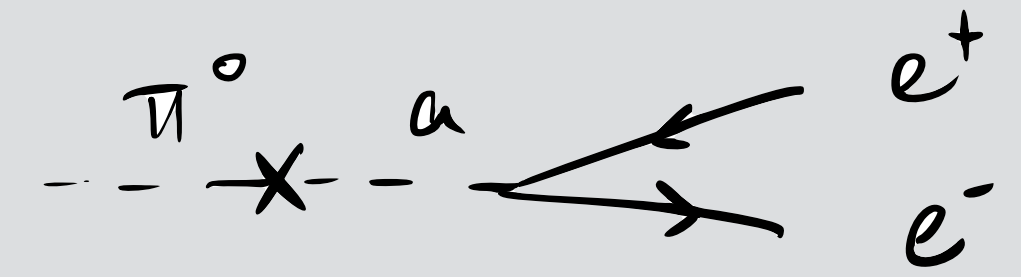
Bertuzzo Jana Machado Zukanovich PRL 2018, PLB 2019  
Arguelles Hostert Tsai PRL 2019  
Ballett Pascoli Ross-Lonergan PRD 2019  
Ballett Hostert Pascoli PRD 2020

## Transition Magnetic Moment



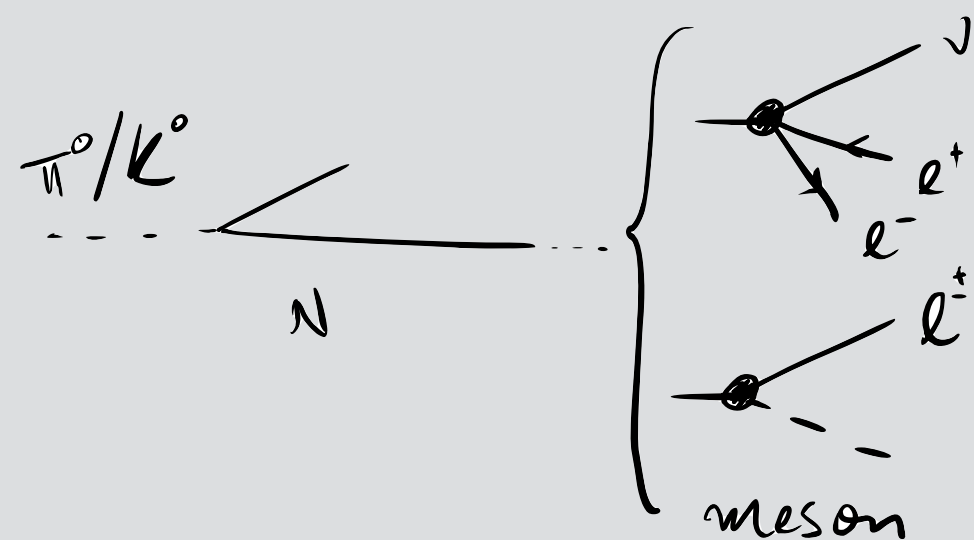
Gninenko PRL 2009  
Coloma Machado Soler Shoemaker PRL 2017  
Atkinson et al 2021 Vergani et al 2021

## Axion-like Particles



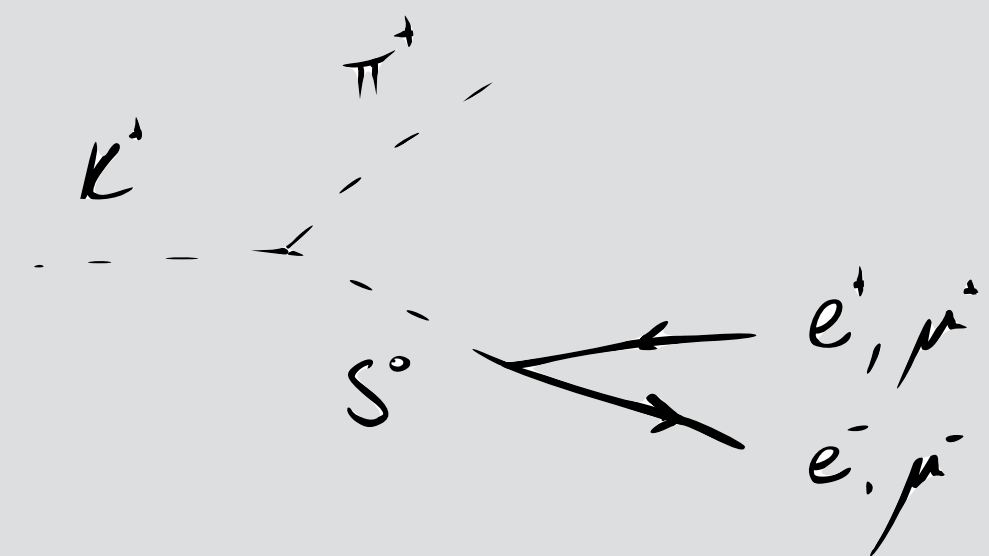
Kelly Kumar Liu PRD 2021  
Brdar et al PRL 2021

## Heavy Neutral Leptons



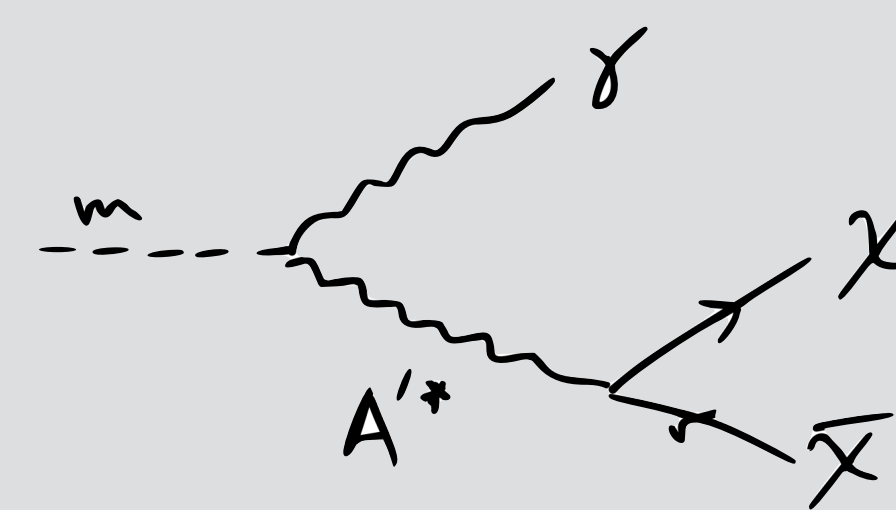
Ballett Pascoli Ross-Lonergan JHEP 2017  
Kelly Machado PRD 2021

## Higgs Portal Scalar



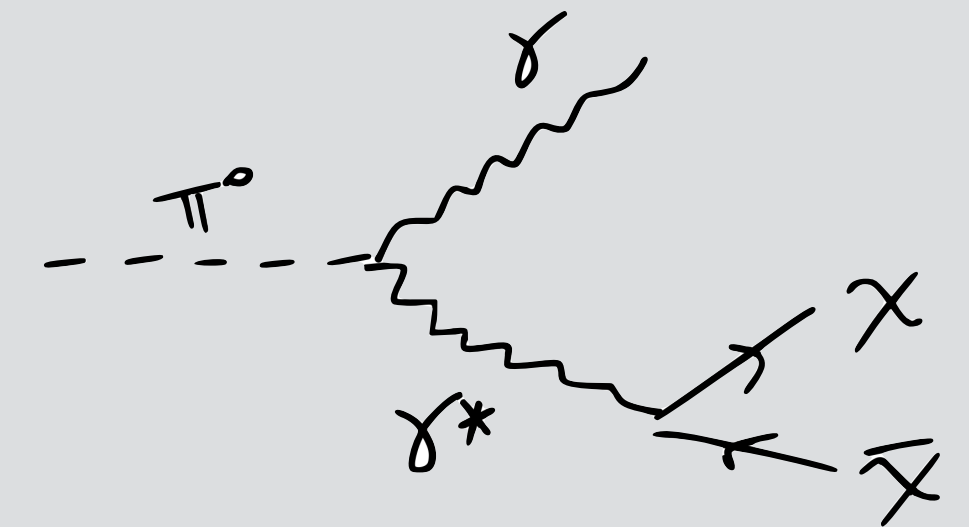
Pat Wilczek 2006  
Batell Berger Ismail PRD 2019  
MicroBooNE 2021

## Light Dark Matter



Romeri Kelley Machado PRD 2019

## Millicharged Particles



Magill, Plestid, Pospelov, Tsai, PRL 2019  
Harnik Liu Palamara, JHEP 2019

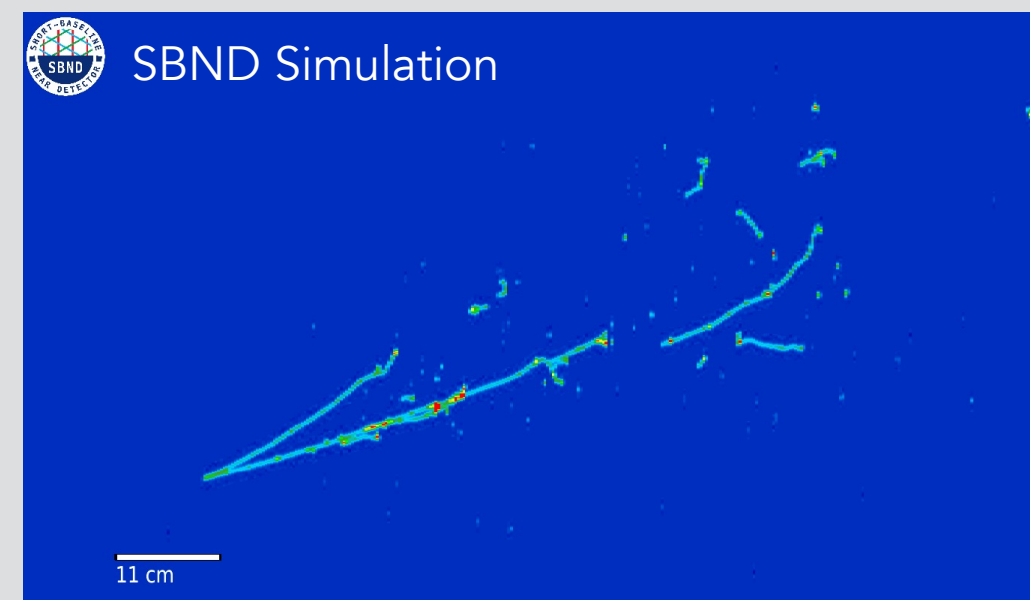
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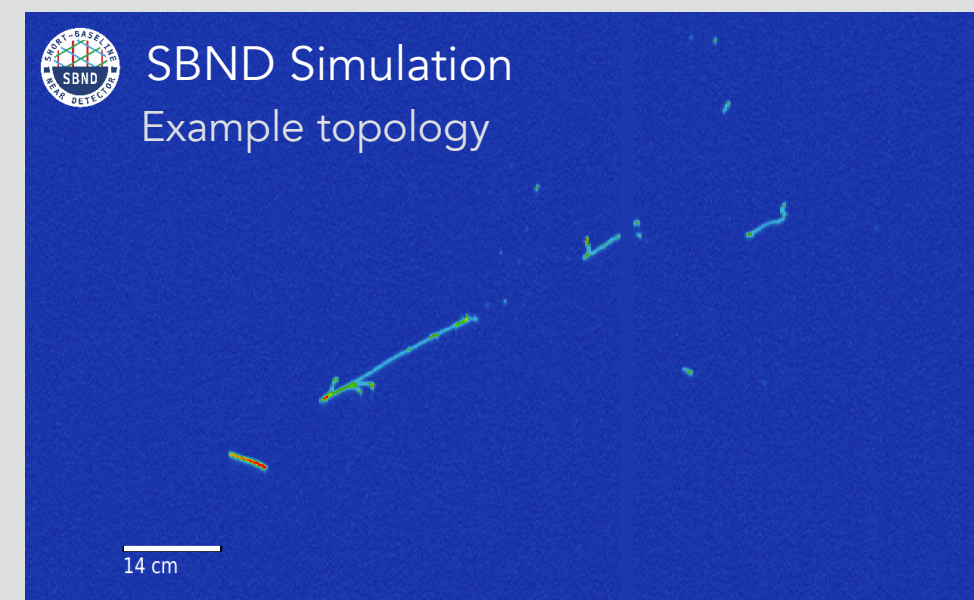
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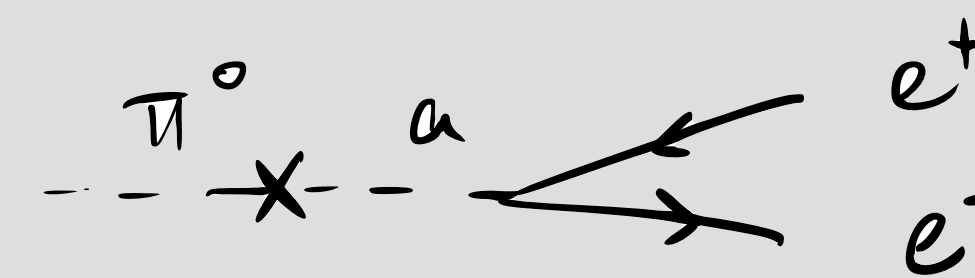
$e^+e^-$  pair w/ or w/o  
hadronic activity

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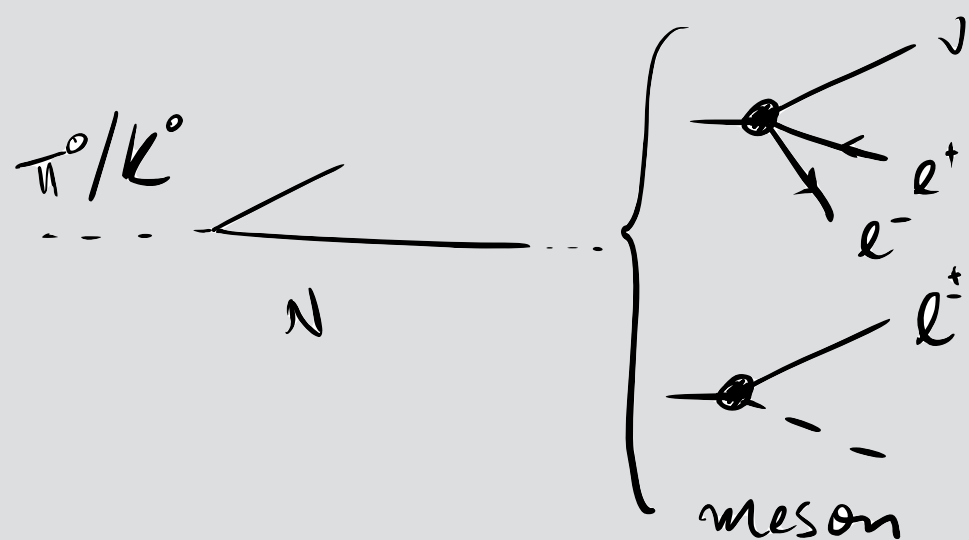
photon shower and  
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## Axion-like Particles



Kelly Kumar Liu PRD 2021  
Brdar et al PRL 2021

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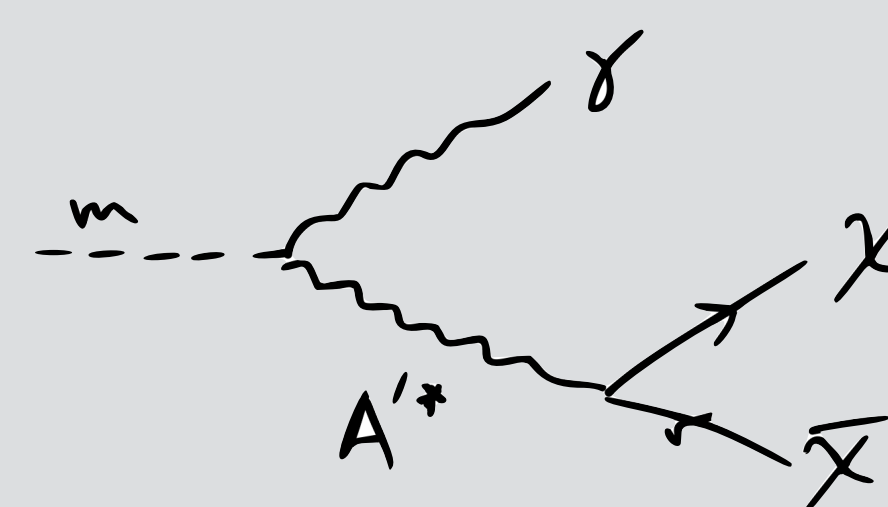
Ballett Pascoli Ross-Lonergan JHEP 2017  
Kelly Machado PRD 2021

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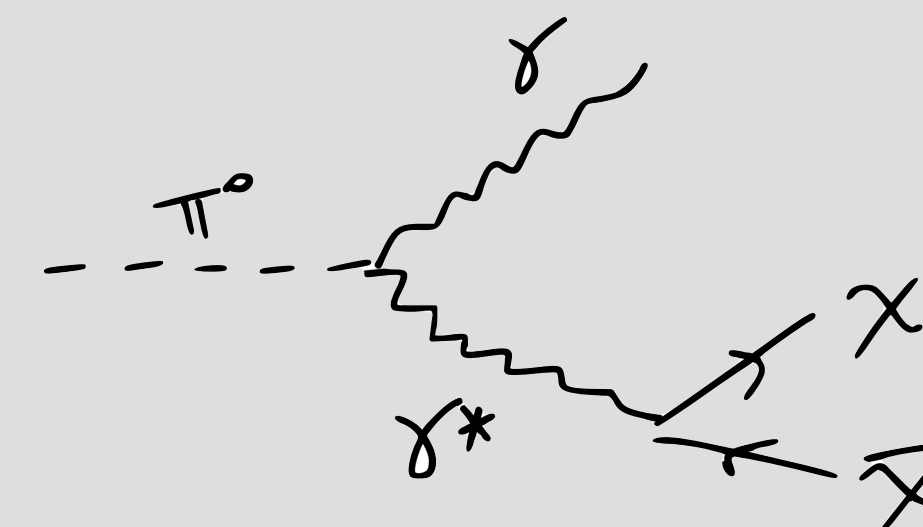
Pat Wilczek 2006  
Batell Berger Ismail PRD 2019  
MicroBooNE 2021

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Romeri Kelley Machado PRD 2019

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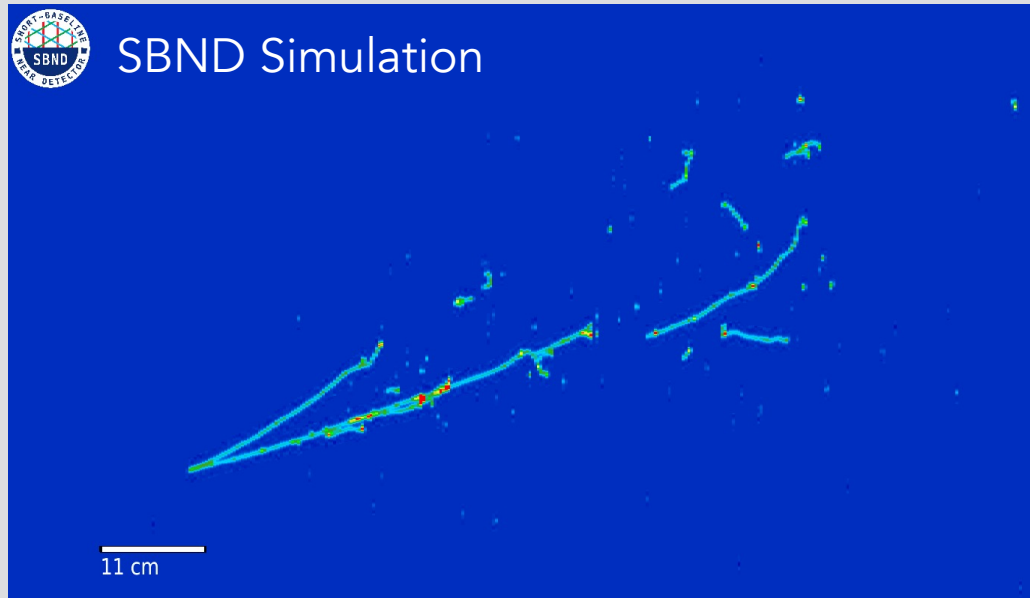
Magill, Plestid, Pospelov, Tsai, PRL 2019  
Harnik Liu Palamara, JHEP 2019

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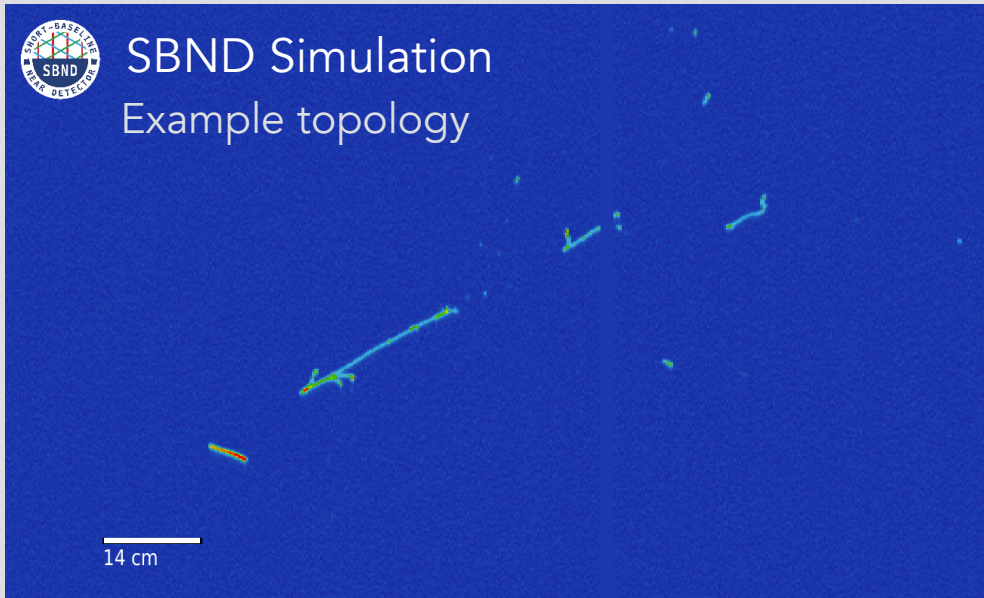
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## Dark Neutrinos



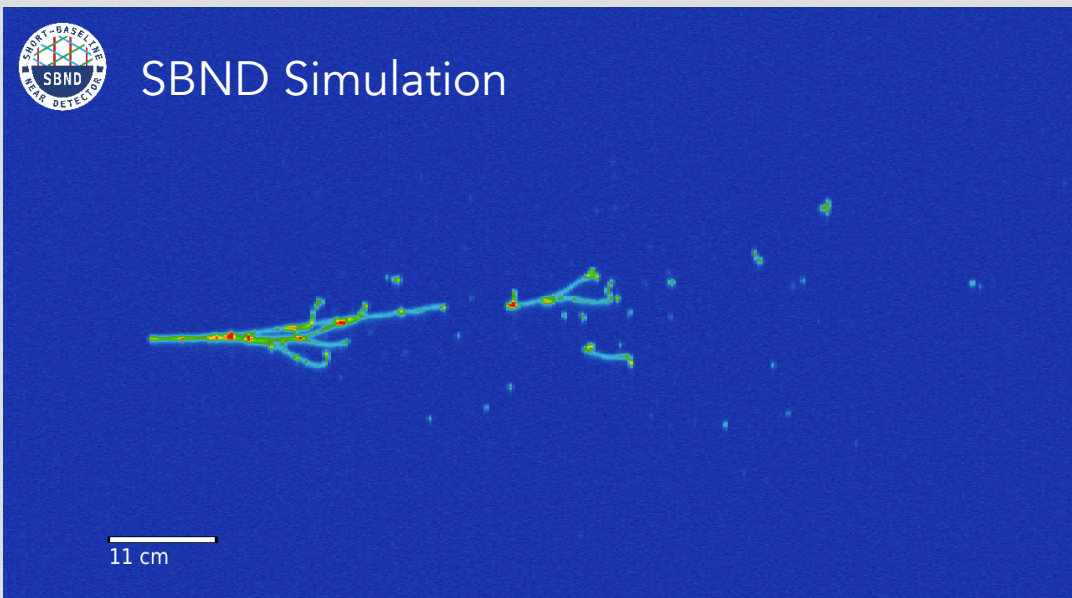
$e^+e^-$  pair w/ or w/o  
hadronic activity

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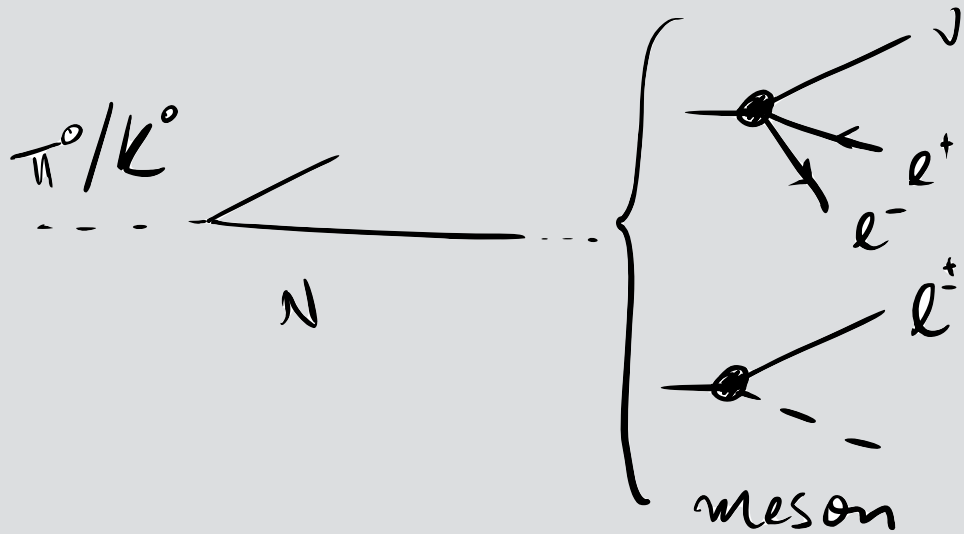
photon shower and  
hadronic activity

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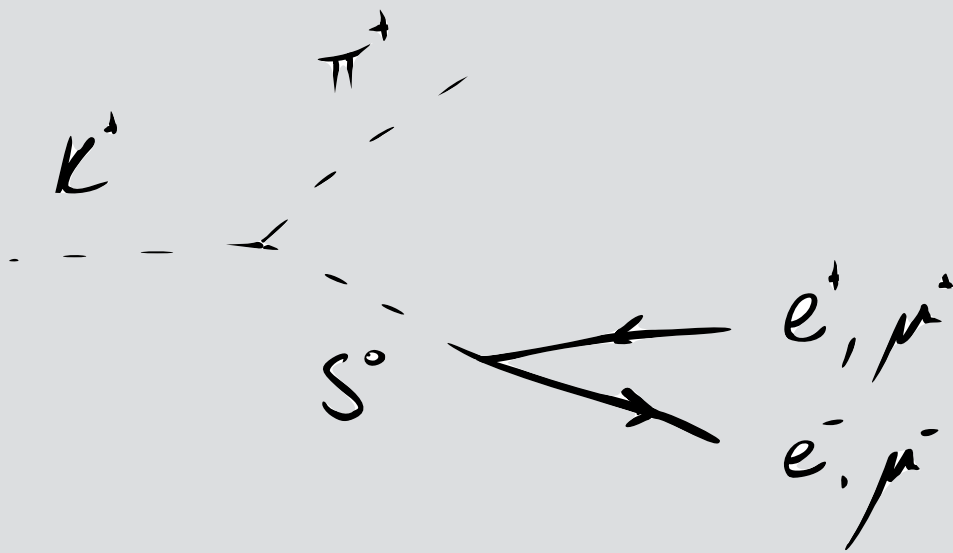
high-energy  
 $e^+e^-$ ,  $\mu^+\mu^-$

## Heavy Neutral Leptons



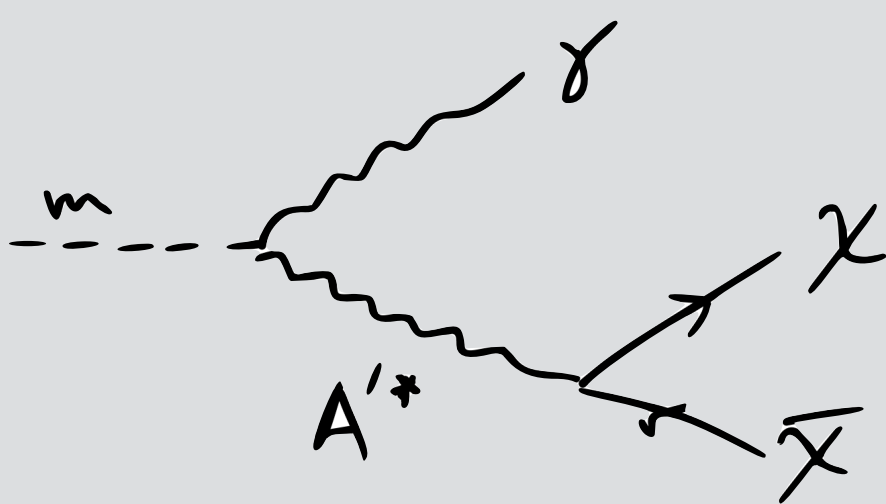
Ballett Pascoli Ross-Lonergan JHEP 2017  
Kelly Machado PRD 2021

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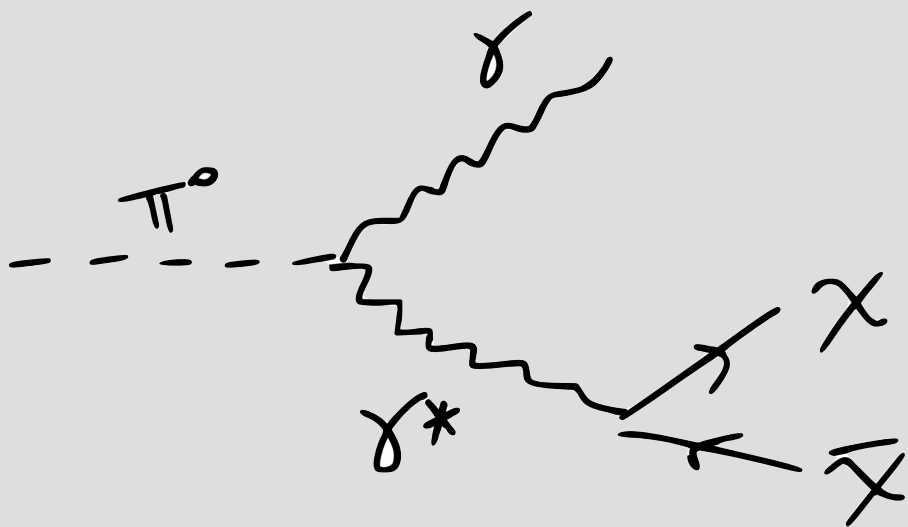
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Batell Berger Ismail PRD 2019  
MicroBooNE 2021

## Light Dark Matter



Romeri Kelley Machado PRD 2019

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Magill, Plestid, Pospelov, Tsai, PRL 2019  
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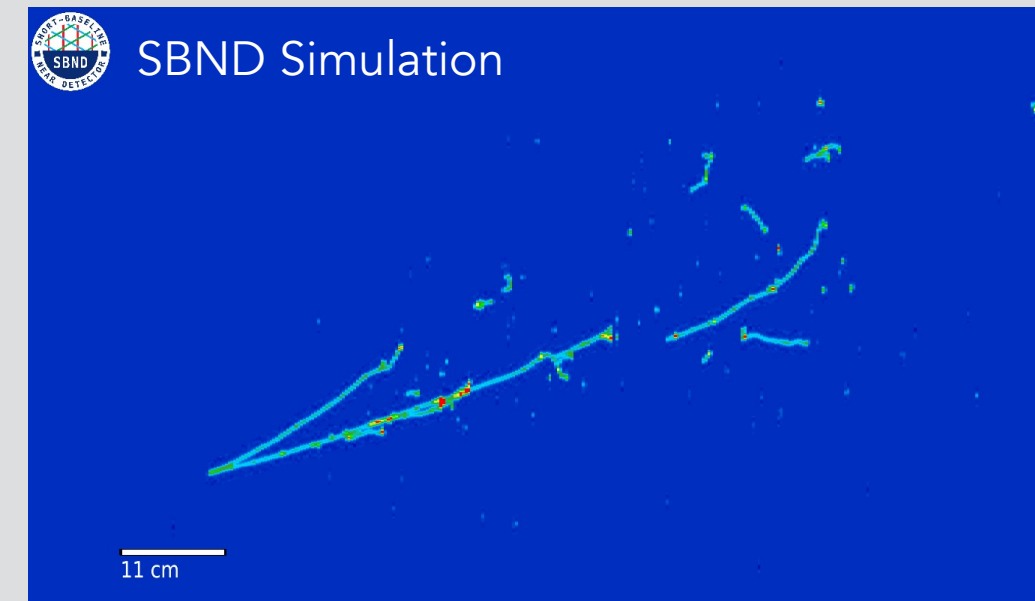
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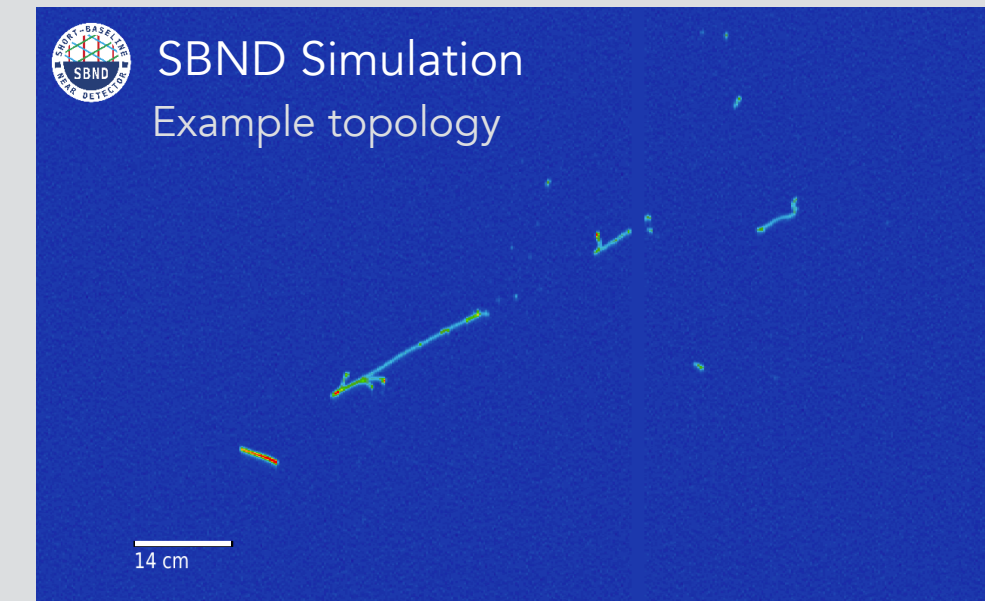
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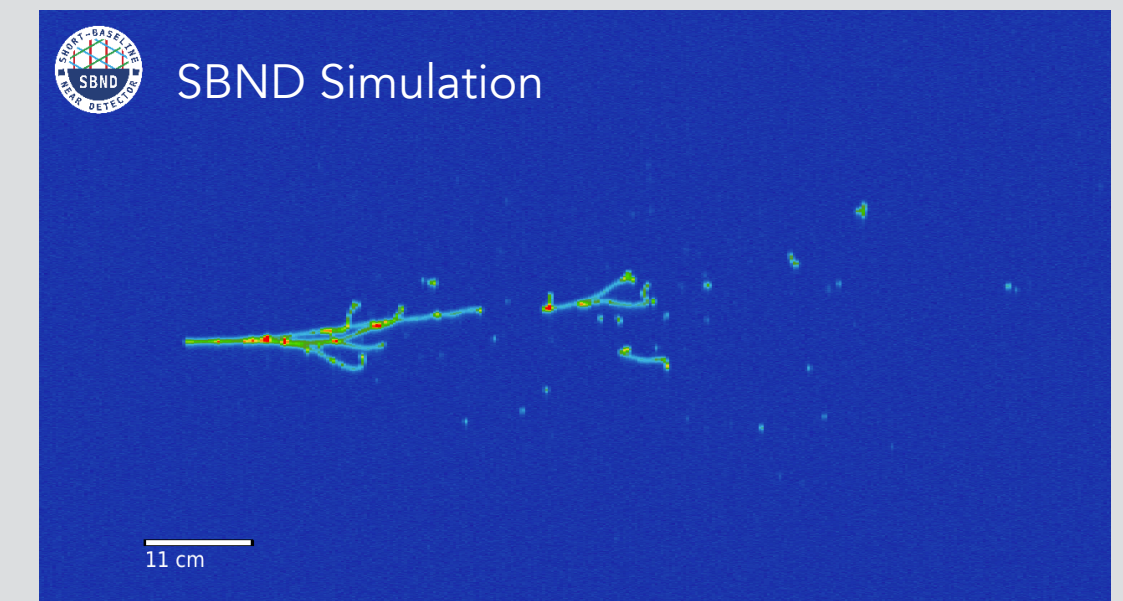
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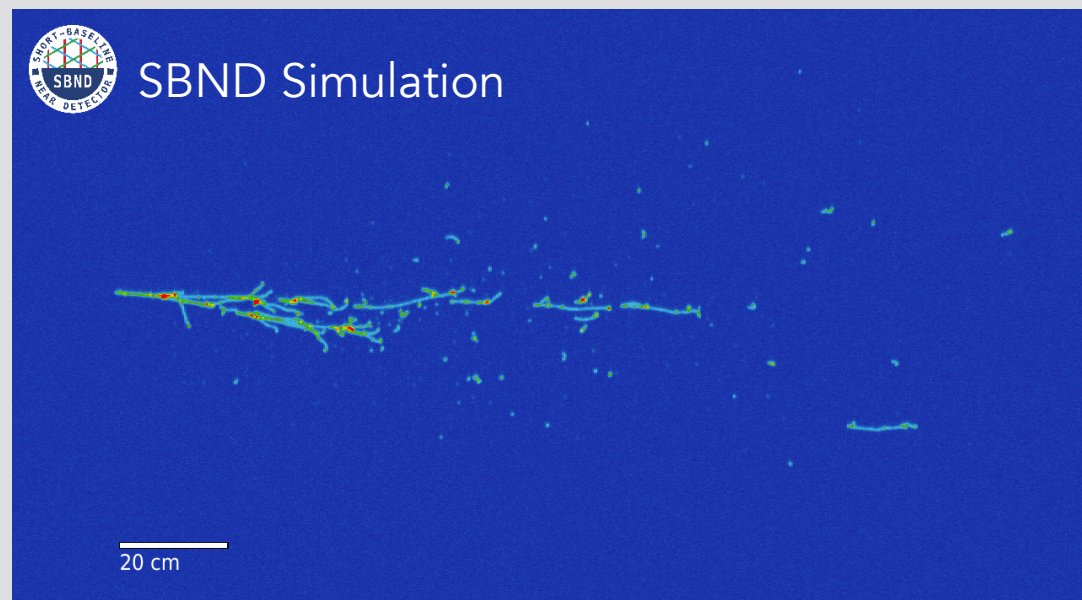
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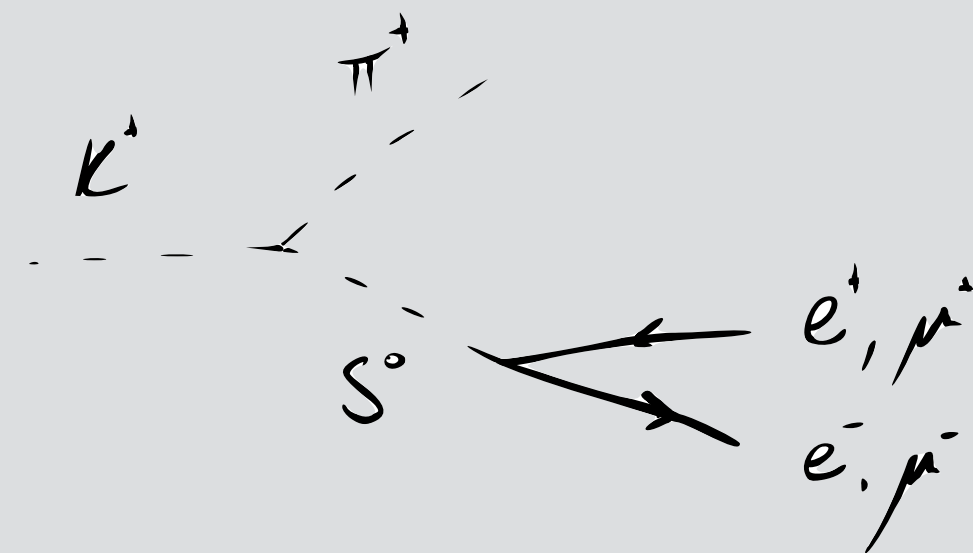
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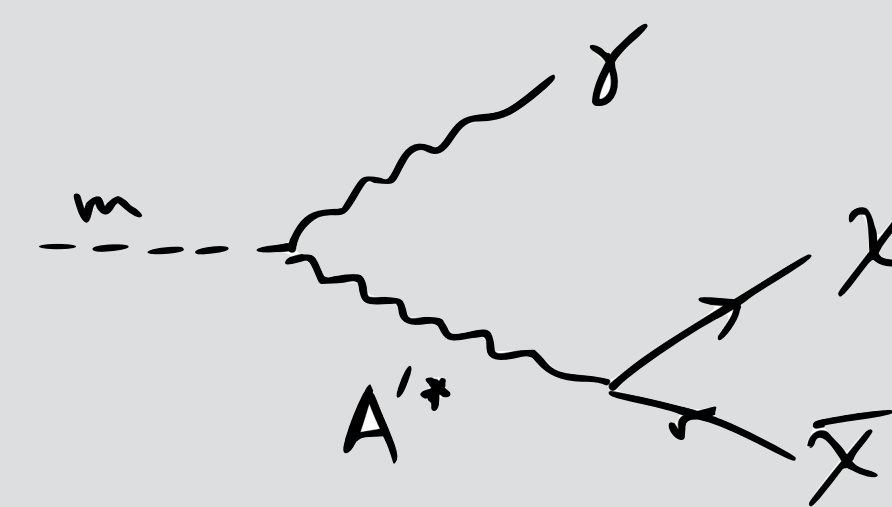
$e^+e^-$ ,  $\mu^+\mu^-$ ,  $\mu\pi$

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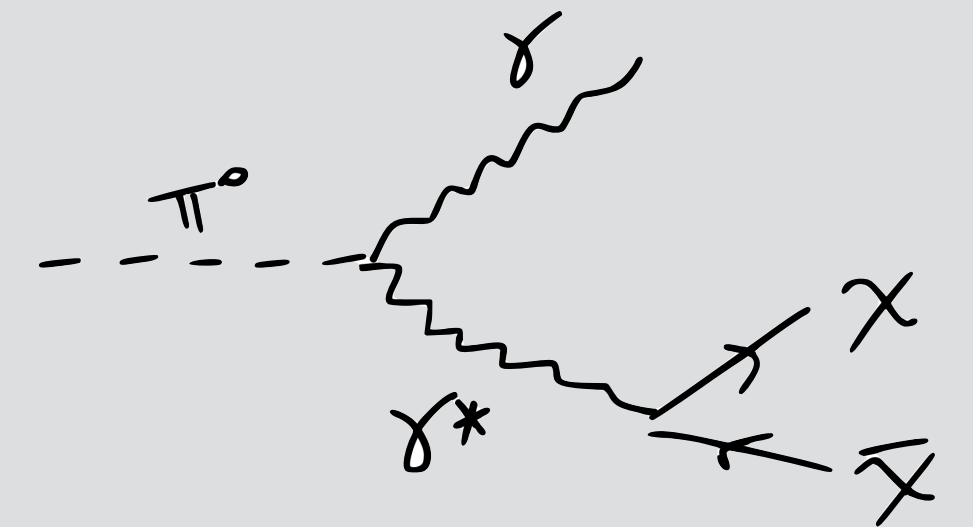
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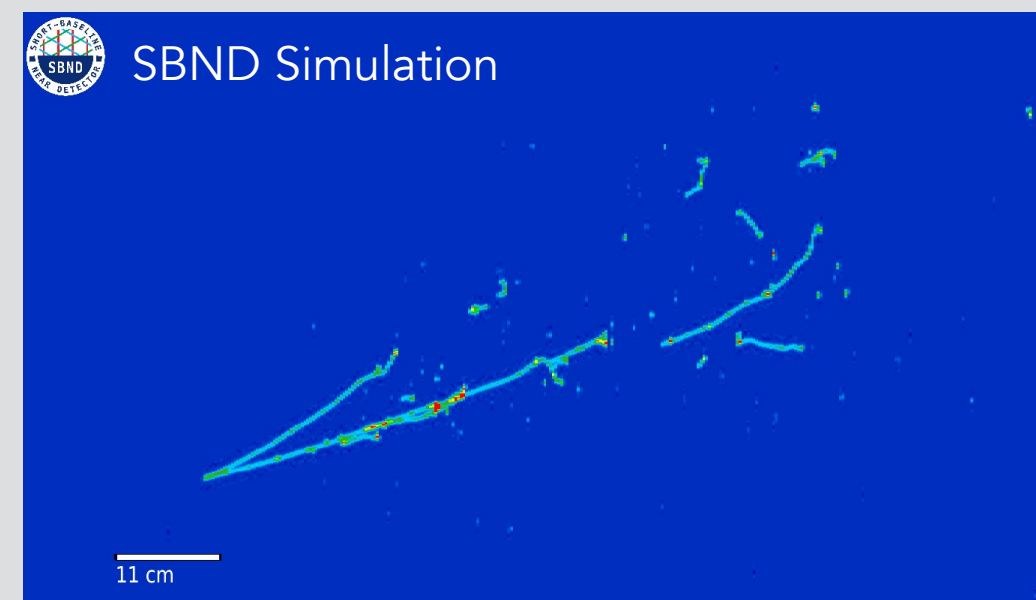
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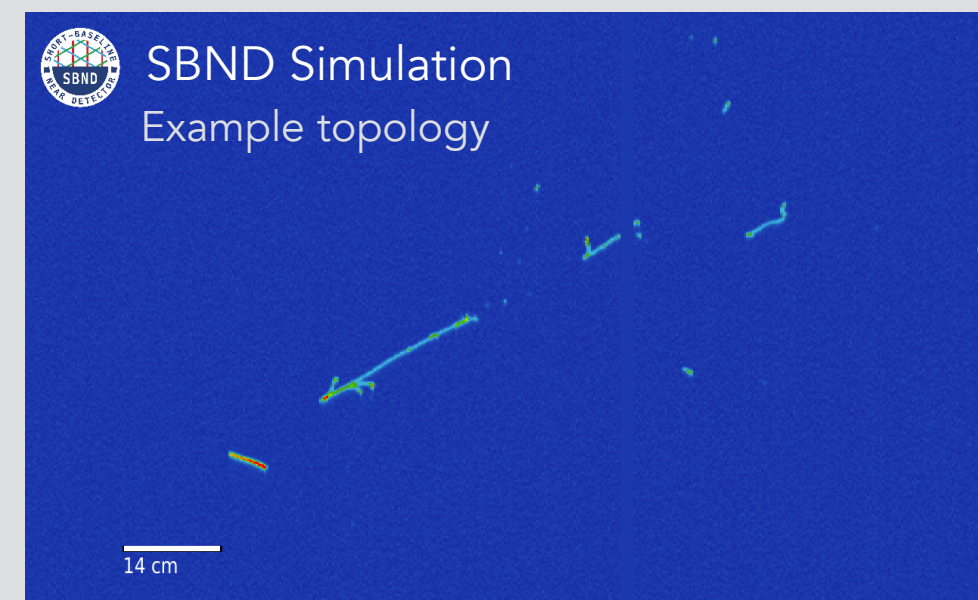
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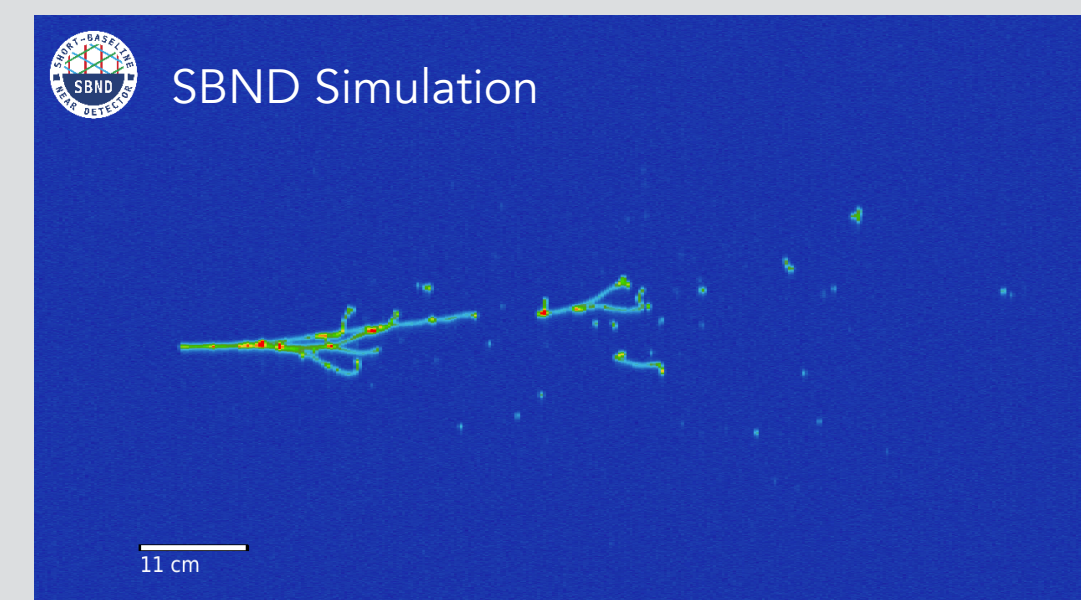
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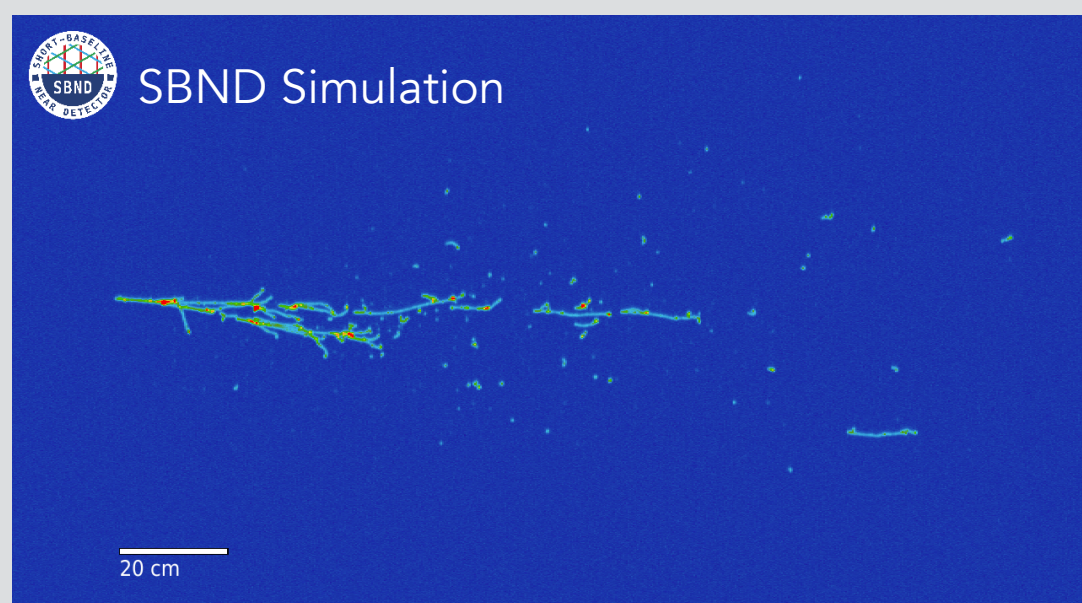
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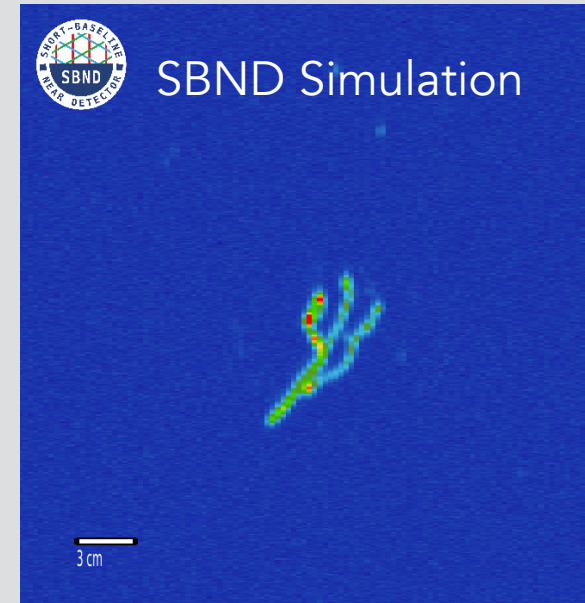
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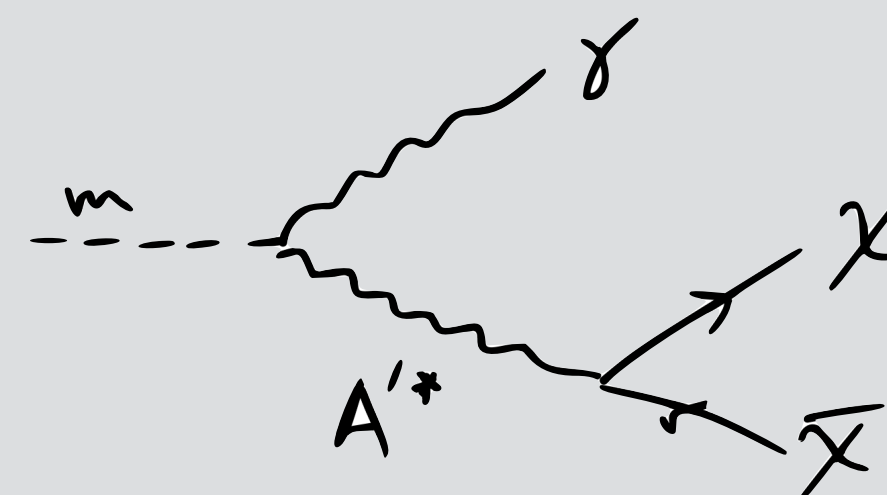
$e^+e^-$ ,  $\mu^+\mu^-$ ,  $\mu\pi$

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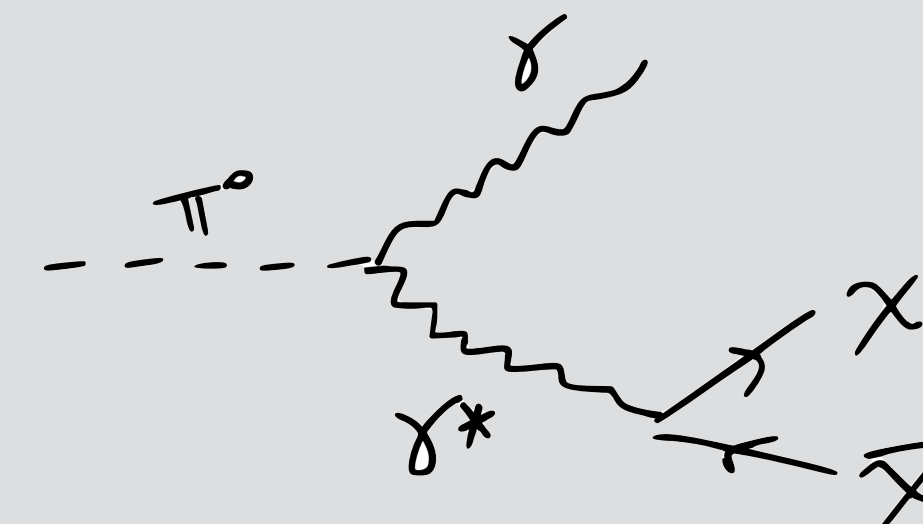
$e^+e^-$ ,  $\mu^+\mu^-$ , no  
hadronic activity

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Romeri Kelley Machado PRD 2019

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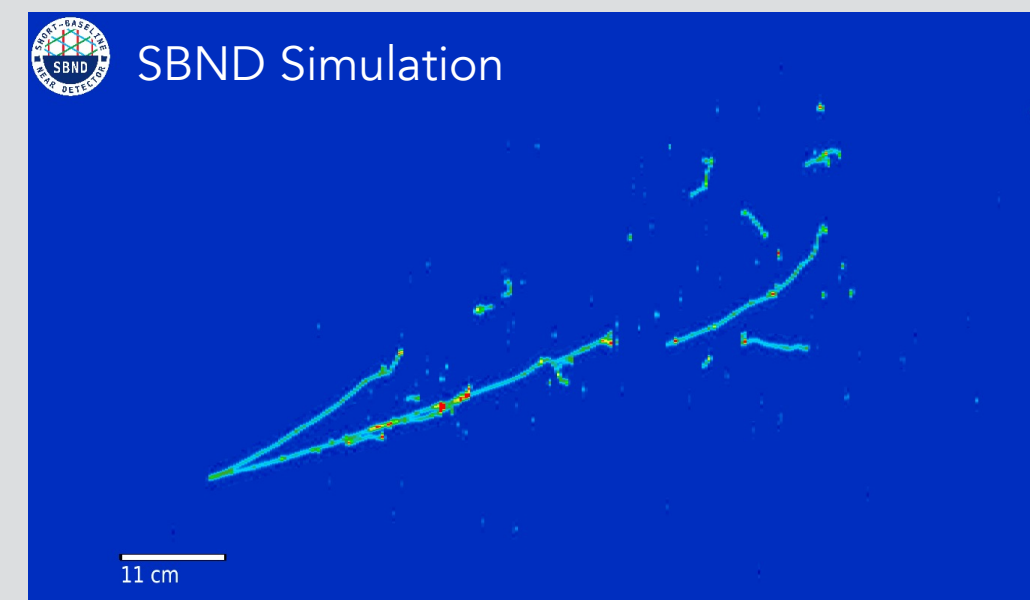
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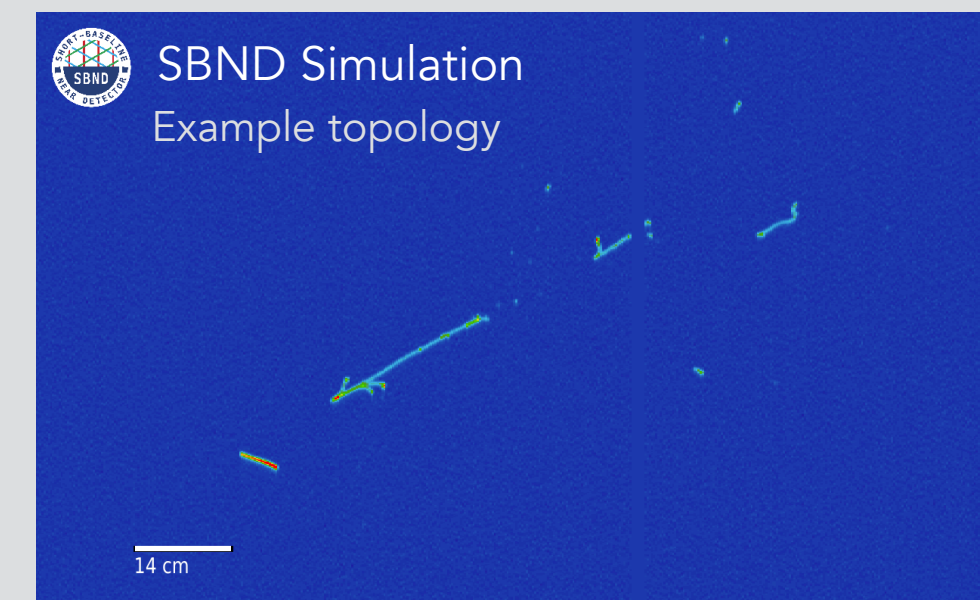
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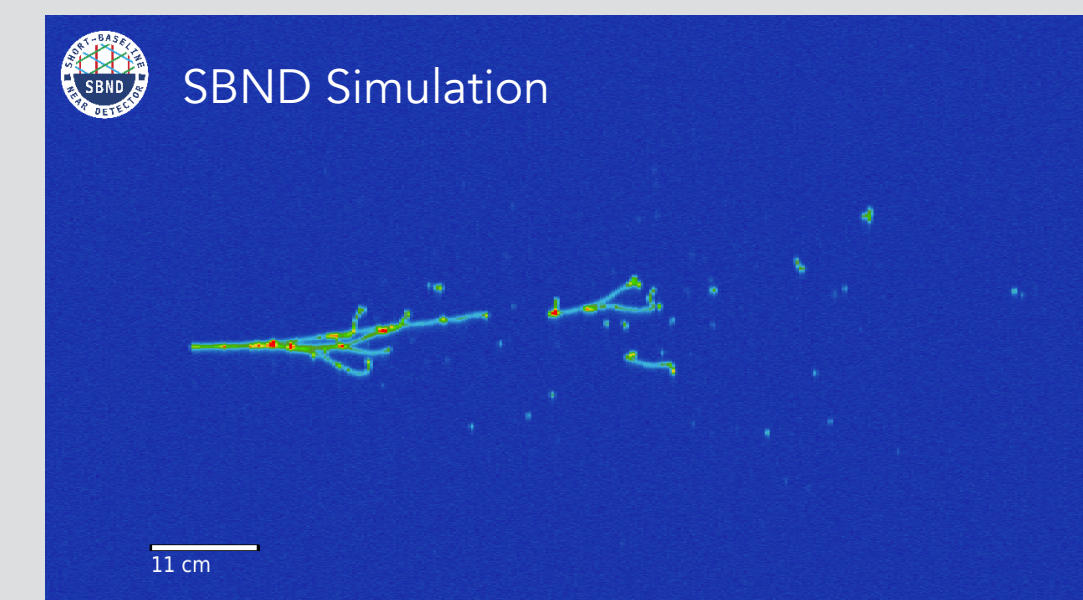
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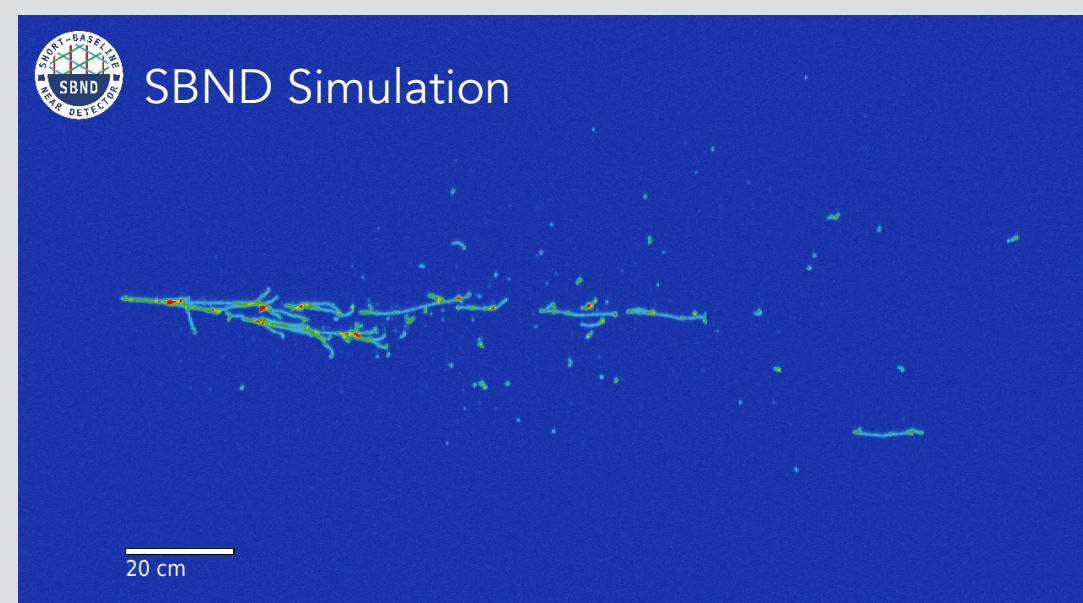
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hadronic activity

## Axion-like Particles



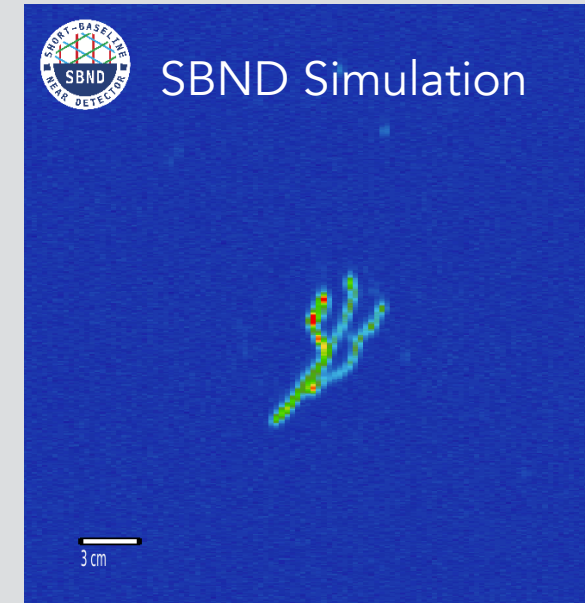
high-energy  
 $e^+e^-$ ,  $\mu^+\mu^-$

## Heavy Neutral Leptons



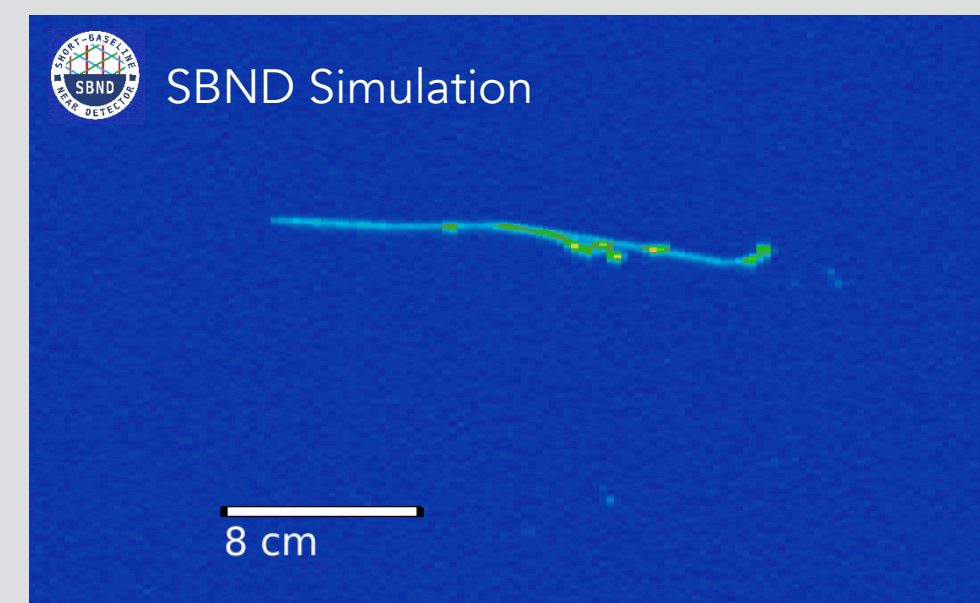
$e^+e^-$ ,  $\mu^+\mu^-$ ,  $\mu\pi$

## Higgs Portal Scalar



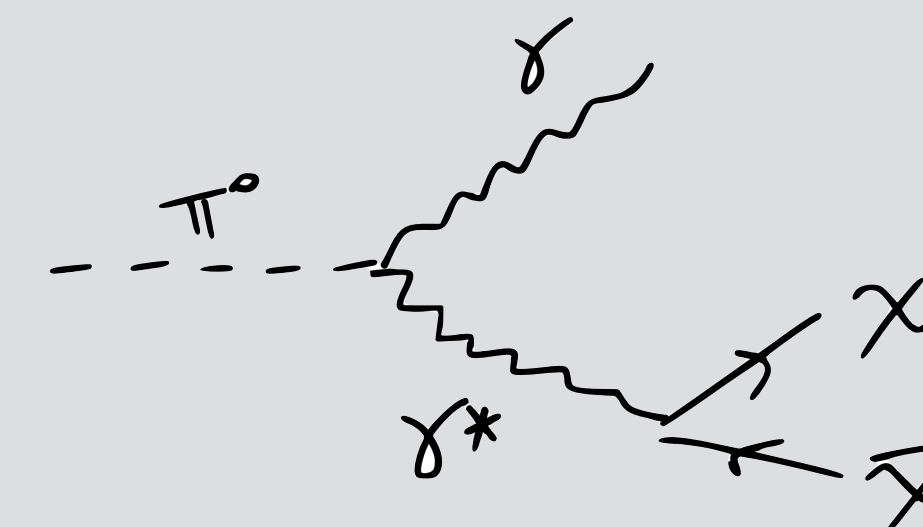
$e^+e^-$ ,  $\mu^+\mu^-$ , no  
hadronic activity

## Light Dark Matter



electron scattering

## Millicharged Particles



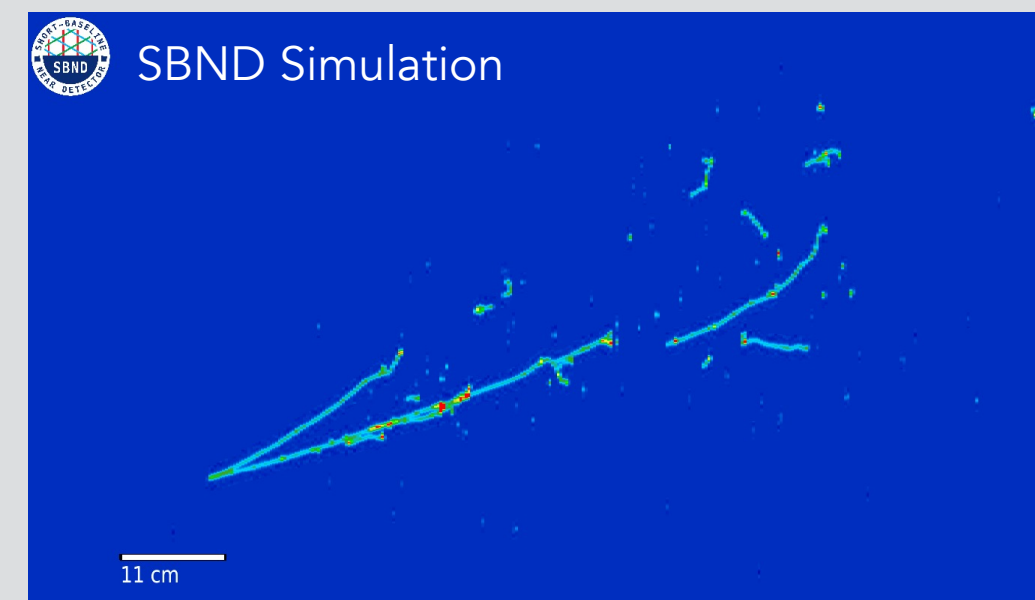
Magill, Plestid, Pospelov, Tsai, PRL 2019  
Harnik Liu Palamara, JHEP 2019

# Beyond Standard Model Searches

Alternative explanations  
to the MiniBooNE excess  
and other BSM scenarios

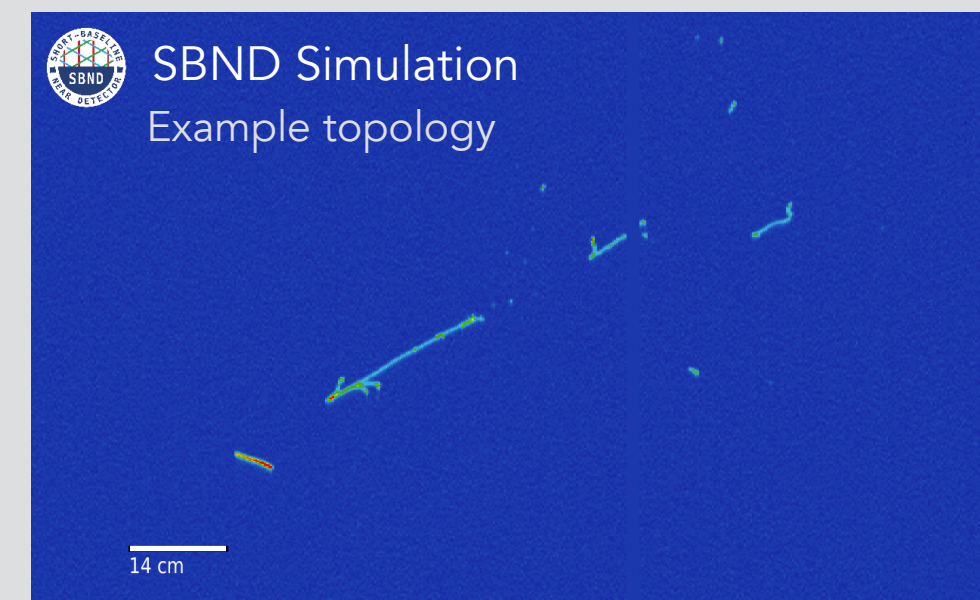
Not an exhaustive list

## Dark Neutrinos



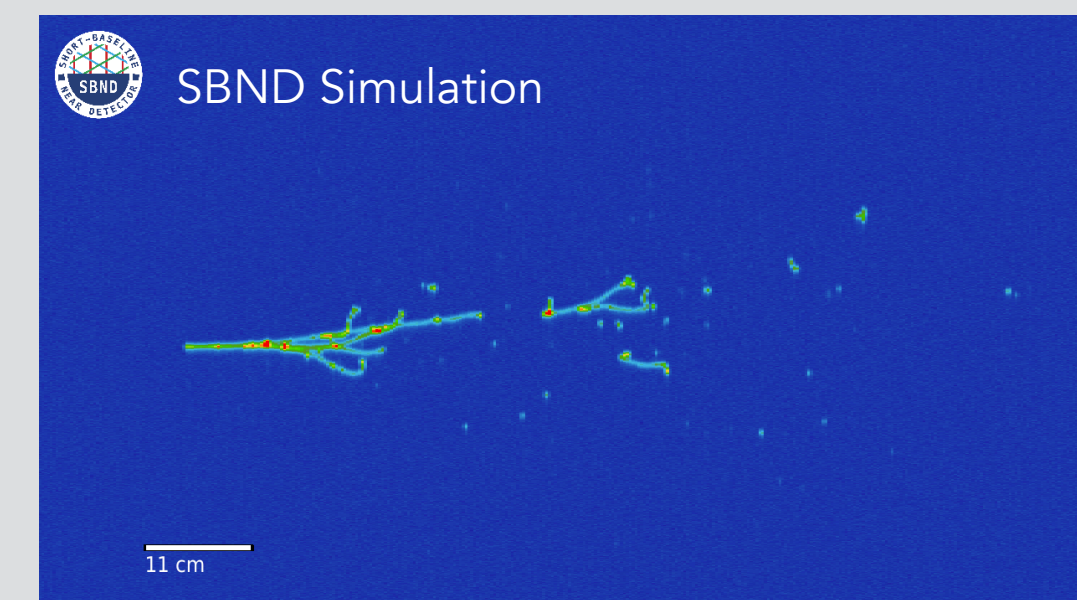
$e^+e^-$  pair w/ or w/o  
hadronic activity

## Transition Magnetic Moment



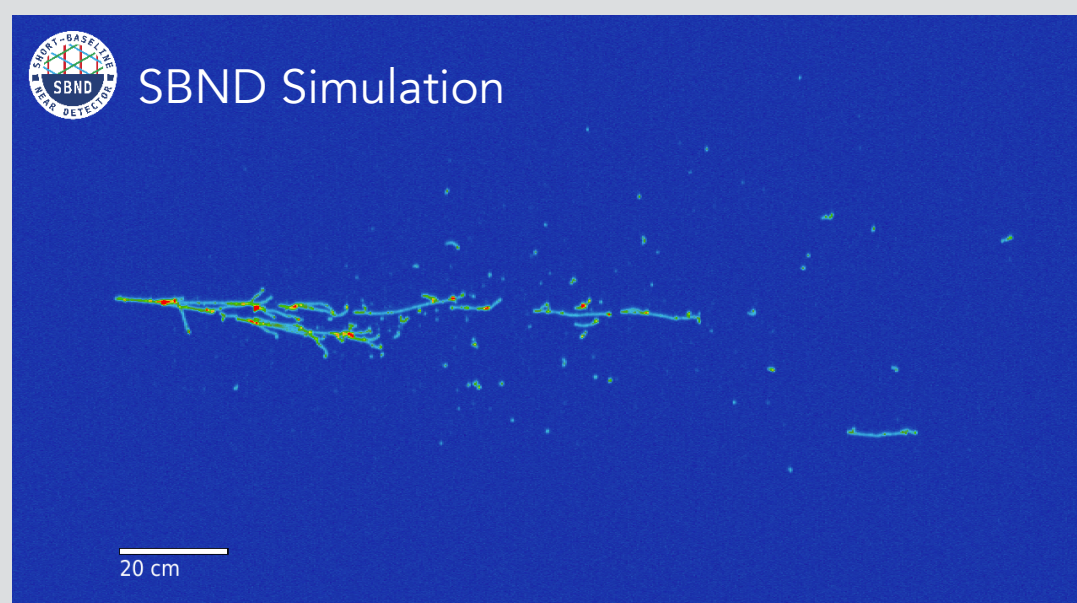
photon shower and  
hadronic activity

## Axion-like Particles



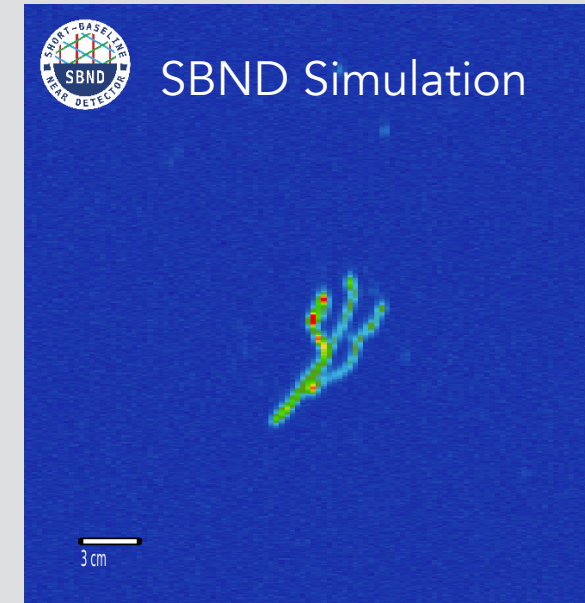
high-energy  
 $e^+e^-$ ,  $\mu^+\mu^-$

## Heavy Neutral Leptons



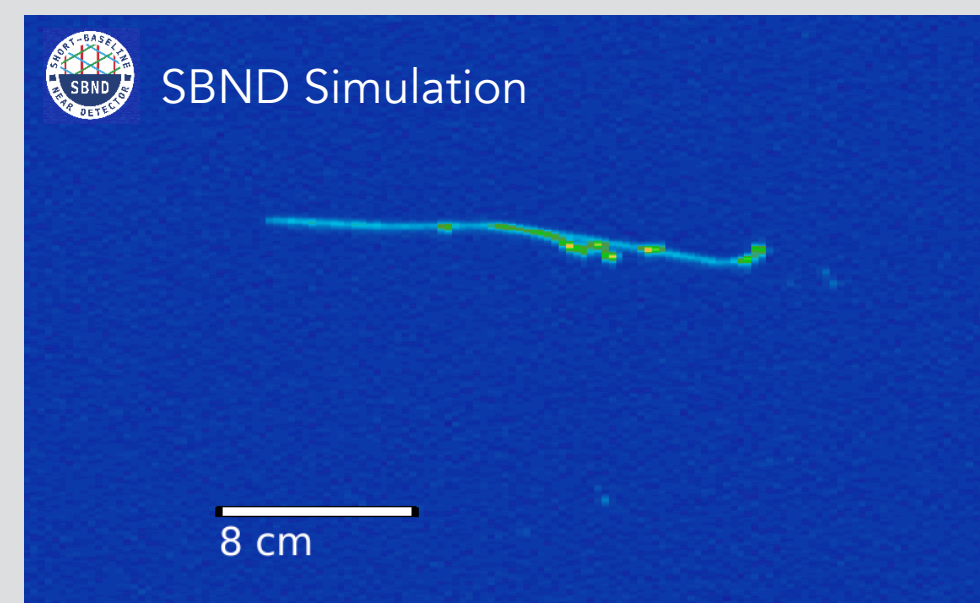
$e^+e^-$ ,  $\mu^+\mu^-$ ,  $\mu\pi$

## Higgs Portal Scalar



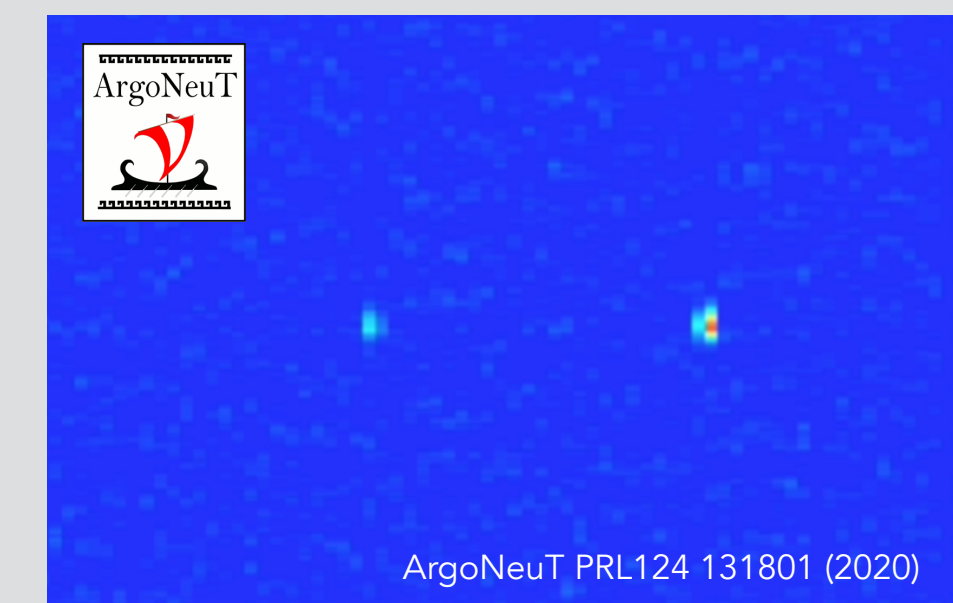
$e^+e^-$ ,  $\mu^+\mu^-$ , no  
hadronic activity

## Light Dark Matter



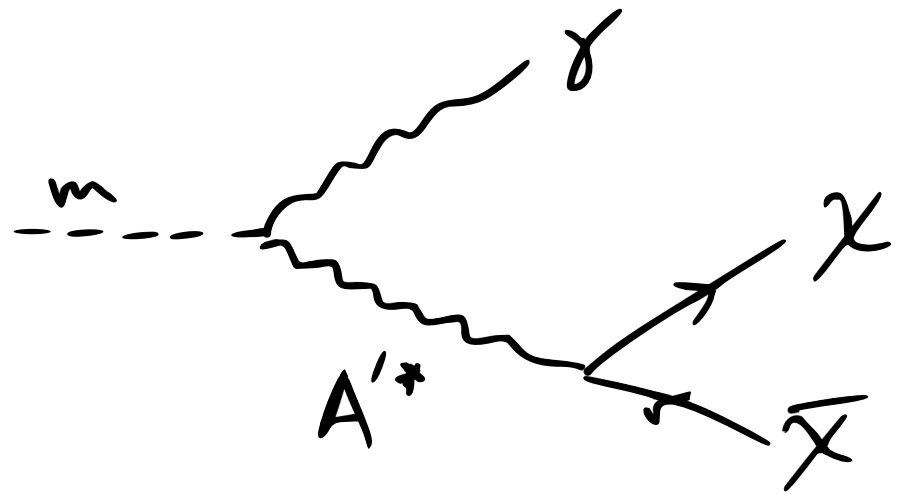
electron scattering

## Millicharged Particles

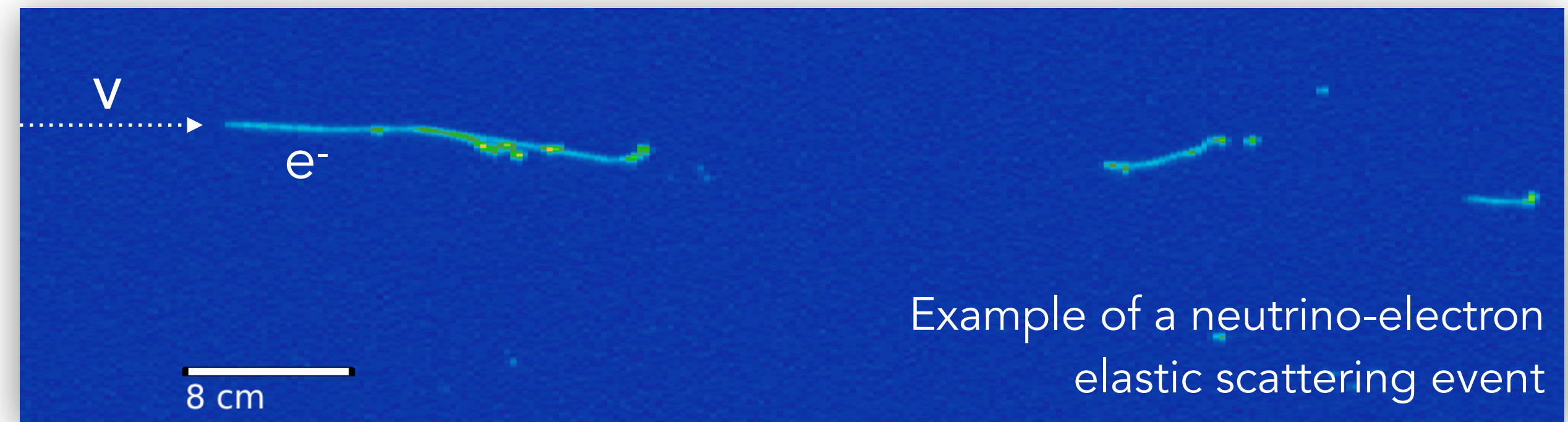
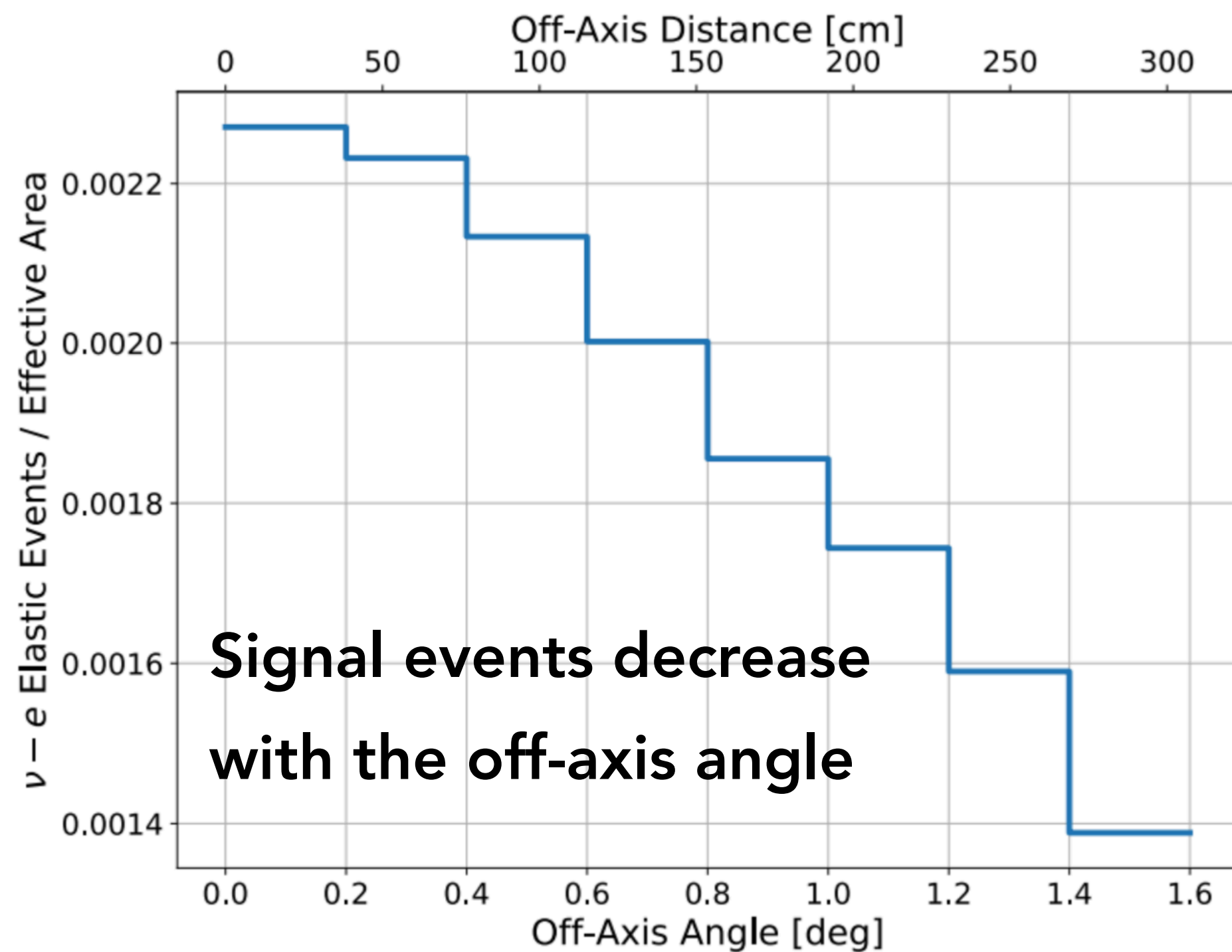


blips/faint tracks

# Dark Matter Searches: Light Dark Matter



One example: **light dark matter** (sub-GeV) coupled to the Standard Model via a dark photon. The dark photons can be produced by neutral meson decays (pions, etas) in the target, and then decay to dark matter. The dark matter can then travel to SBND and, through the dark photon, **scatter off electrons in the detector**.



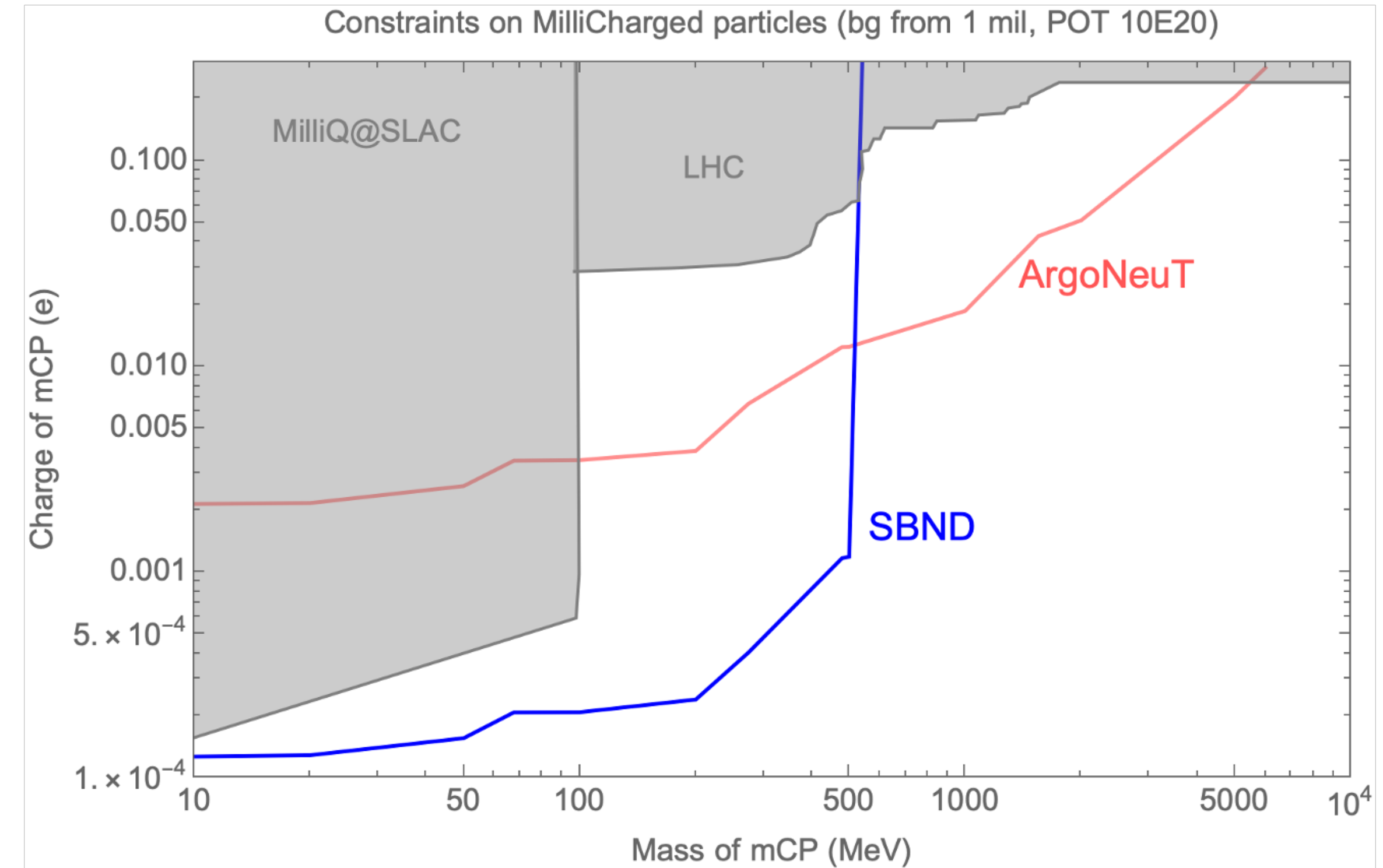
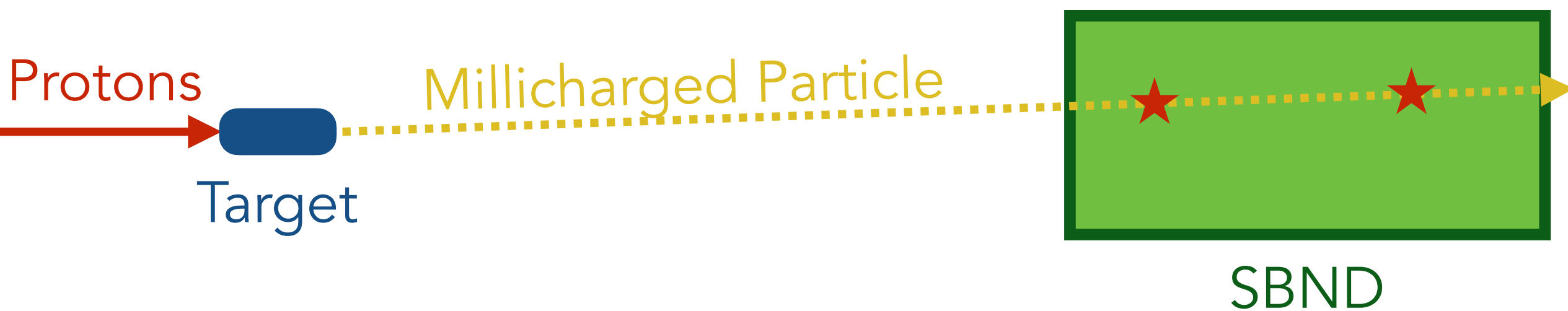
- **Background:** neutrino-electron elastic scattering. Neutrinos come from two-body decays of charged (focused) mesons.
- **Signal:** elastic scattering electron events. Dark matter comes from three-body decays of neutral (unfocused) mesons.
- **Neutrino flux drops off more sharply as a function of radius!**

See also <https://arxiv.org/abs/1903.10505>

# Dark Matter Searches: Millicharged Particles

- **Millicharged particles:** hypothetical new particles with fractional charge.
- Neutral mesons produced from proton collisions with the target could decay into millicharged particles.
- Millicharged particles will produce low-energy depositions (small hits or faint tracks) that point back to the target.
- **SBND could provide a promising new search for millicharged particles.**

Argoneut method: R. Acciarri et al., PRL124 131801 (2020)



Preliminary results from simulation

# Conclusions

- SBND detector construction and installation are progressing very well
- We expect to begin detector operations next year
- SBND will have an extensive physics program including:
  - search for eV mass-scale sterile neutrinos oscillations in the SBN program
  - neutrino cross-section measurements
  - search for new and exotic physics signals



**SBND TPC was  
completed this week!**